

## **Introduction:**

- One of the key measurements from the Clouds and the Earth's Radiant Energy System (CERES) satellite and the Baseline Surface Radiation Network (BSRN) is Earth emitted or longwave (LW) radiation.
- The disestablished CERES Ocean Validation Experiment (COVE), located at Chesapeake Light Station was a validation site for CERES measurements and collected upwelling LW (LW<sup> $\uparrow$ </sup>) with a static pyrgeometer, but the measurement was complicated due to the light station tower being in its field of view, which we estimate to be 15%.
- To resolve the tower issue, we obtain a different LW<sup> $\uparrow$ </sup> value (LW<sup> $\uparrow$ </sup><sub>f=0</sub>) using data from other instruments located at COVE such as an Infrared Radiation Thermometer (IRT) to measure ocean skin temperature, downwelling pyrgeometer and meteorological sensors.
- Comparing the static pyrgeometer  $(LW_{pvr}^{\uparrow})$  with  $LW_{f=0}^{\uparrow}$  shows the unwanted consequence of the tower, evident on a clear, summer day. However, winter comparisons are biased even more with differences up to 5% (20 W/m^2). BSRN target uncertainty is 2%.



COVE at Chesapeake Light Station information:

- ~25km off the coast of Southeast Virginia, USA
- Coordinates: 36.90N, 75.71W
- Water depth is ~12m
- Operational from May 1999 December 2016

ummer Clear



pyrgeometers  $(LW_{pvr}^{\uparrow})$  field of view. The height of layer  $\varepsilon_1$  is 21m.



than winter clear but  $LW_{f=0}^{\uparrow}$  values are still distinctly higher.

**Right:** A single year plot of  $LW_{pvr}^{\uparrow}$  minus  $LW_{f=0}^{\uparrow}$  when  $\varepsilon_1 = 0$ . Notice the large differences between the two values as indicated by the color bar (darker reds in late afternoon and dark blues in the winter and nighttime). Ideally, one should see all white. This plot encapsulates the need to find a better LW<sup>†</sup>value.

## Measuring Upwelling Longwave in the Presence of an Obstruction Bryan Fabbri<sup>1</sup>, Greg Schuster<sup>2</sup>, Fred Denn<sup>1</sup>, Bing Lin<sup>2</sup>, Jay Madigan<sup>1</sup>, Robert Arduini<sup>1</sup>

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**Left:** Fisheye lens picture showing the approximate 15% tower obstruction in the static

**Right:** Illustration of the instrument geometry at the Chesapeake Light Station.  $LW_{pvr}^{\uparrow}$ measurements are located at the end of on an 8m boom, which is not long enough to remove the platform from the field of view.  $LW^{\downarrow}$  flux measurements and narrowband IRT have an unobstructed field of view at the top of the tower. The emissivity's  $\varepsilon_{atm}$ and  $\varepsilon_1$  are the atmospheric column and the layer of air below the boom, respectively.

**Left**:  $LW_{pvr}^{\uparrow}$  measurements (black), air temperature converted to

The line plot colors for the three single day scenarios are the same as Summer Clear. a) Summer Overcast. There is not a dramatic difference between  $LW_{pvr}^{\uparrow}$  and  $LW_{f=0}^{\uparrow}$  values, suggesting low temperature differences between the tower and water. **b**) Winter Clear. The disparity between  $LW_{pvr}^{\uparrow}$  and  $LW_{f=0}^{\uparrow}$  are greatest among the four scenarios, implying the tower obstruction plays an even greater role in the wintertime. c) Winter Overcast. Differences between  $LW_{pvr}^{\uparrow}$  and  $LW_{f=0}^{\uparrow}$  are less



Where,

Website: <u>https://science.larc.nasa.gov/CRAVE/</u>

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$$= (1 - \varepsilon_1)[\varepsilon_w \sigma T_w^4 + (1 - \varepsilon_1)(1 - \varepsilon_w)LW^{\downarrow}] + \varepsilon_1 \sigma T_1^4$$

- $\varepsilon_1$  = Emissivity of the air below the static pyrgeometer
- $\varepsilon_{\rm w}$  = Emissivity of the water (0.99)
- $\sigma$  = Stefan-Boltzmann constant (5.67 x 10<sup>-8</sup>)
- $\Gamma_{w} = Water temperature (°K)$
- $_{1}$  = Air temperature (°K)
- $LW^{\downarrow}$  = Downwelling LW radiation

Left: The new method to determine LW<sup>†</sup> without tower interference. This equation includes water emission ( $\varepsilon_w \sigma T_w^4$ ), water reflectance after attenuation by the air below our sensors  $[(1 - \varepsilon_1)(1 - \varepsilon_w)LW^{\downarrow}]$ , and emission of the air below our sensors  $(\varepsilon_1 \sigma T_1^4)$ . It also accounts for LW attenuation from the surface to the sensor. **Right**: Determining  $\varepsilon_1$ . Since 1 -  $\varepsilon_1$  is the transmissivity of the layer below the boom, we can relate this to the longwave optical depth,  $\tau_1$ . Since nearly all the LW absorption is attributable to water vapor; we can scale the optical depth to the fraction of water vapor in  $\varepsilon_1$ . We have everything we need to obtain  $\varepsilon_1$  except the emissivity of the atmosphere ( $\varepsilon_{atm}$ ). Since  $\varepsilon_{atm}$  is not precisely known and varies with temperature and dewpoint, we estimate the range of  $\varepsilon_1$  using the full range of values of  $\varepsilon_{atm}$  (0-0.99). Inserting  $\varepsilon_{atm}$  values and propagating the resulting  $\tau_1$  indicates  $\varepsilon_1$  ranges from 0.0 to 0.049. These values are used to calculate  $LW_{f=0}^{\uparrow}$ , which accounts for the air and water below the boom but does not include the effects of the tower.

Note: The orange plot lines in the lower left of this poster are the full range of  $\varepsilon_1$ . Dark blue is  $\varepsilon_1 = 0$ .





## Summary:

- 15% obstruction in the pyrgeometer field of view, first noticeable on a clear, summer day.
- the BSRN 2% target uncertainty and point to using  $LW_{f=0}^{\uparrow}$  as an improved measurement.
- field of view issues.





$$(1 - \varepsilon_1) = exp(-\tau_1)$$
  
$$\tau_1 = -\left[\frac{Q_1\rho_1Z_1}{W}\right] \ln(1 - \varepsilon_{atm})$$

Where,

 $Q_1$  = Water vapor mixing ratio of layer  $\varepsilon_1$ (Obtained from meteorological data)  $\rho_1$  = Density of dry air of layer  $\varepsilon_1$  (1.225 kg/m<sup>3</sup>)  $Z_1$  = Height of layer  $\varepsilon_1$  (21m) W = Column precipitable water vapor  $\varepsilon_{\rm atm}$  = Emissivity of the atmosphere

COVE's estimated tower obstruction is 15% (F = 0.15). Approximately one-third of the relative bias is outside the target uncertainty when F = 0.15. Therefore, the solution to the tower obstruction is to use the output determined from  $LW_{f=0}^{\uparrow}$  while also meeting BSRN target uncertainty. If we were able to move the  $LW_{pvr}^{\uparrow}$  instrument closer to the tower (when F gets larger), the results get worse. If we were able to move the  $LW_{pvr}^{\uparrow}$  instrument further away from the tower (when F gets smaller), the numbers will eventually match  $LW_{f=0}^{\uparrow}$ . At COVE, the boom would need to extend to ~14m to be inside the target uncertainty at F = 0.05.

• Many years of LW<sup>↑</sup> data collected with a static pyrgeometer at COVE were contaminated with an estimated

• A new LW<sup>†</sup> value (LW<sup>†</sup><sub>f=0</sub>) was determined and the only unknown from the equation was air emissivity ( $\varepsilon_1$ ) below the static pyrgeometer measurement. Applying the full range of  $\varepsilon_1$  to  $LW_{f=0}^{\uparrow}$  display small differences. • Comparing  $LW_{f=0}^{\uparrow}$  with  $LW_{pvr}^{\uparrow}$  shows noticeable differences, more pronounced in the winter, that are outside

• Using  $LW_{f=0}^{\uparrow}$  makes the tower obstruction obsolete. This method could be used at other locations with similar

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