

High image resolution NASA CALIOP extinction denoising / inference

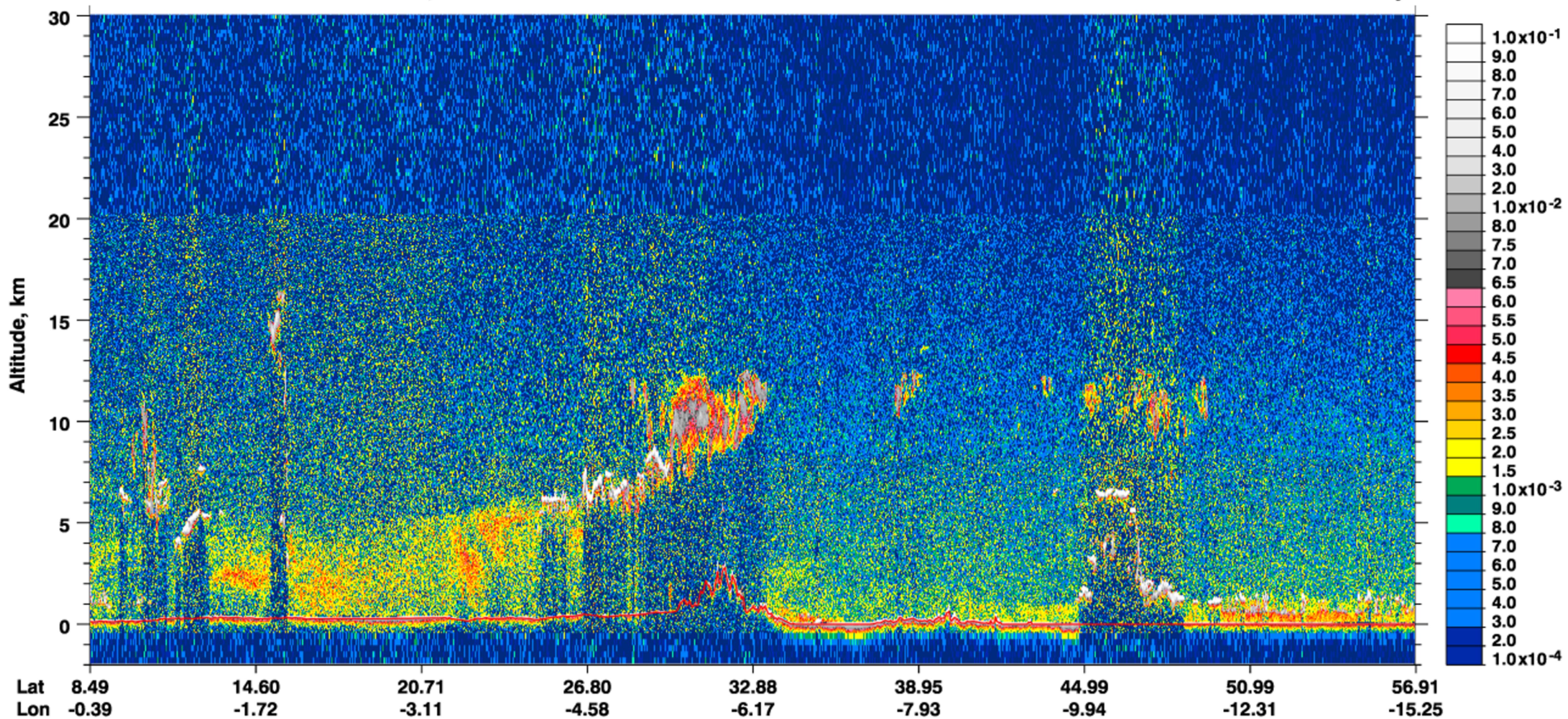
Constructive insights for future space-based missions

**Willem J. Marais, Robert E. Holz, Mark A. Vaughan, Charles R. Trepte,
John W. Hair, Chris A. Hostetler**

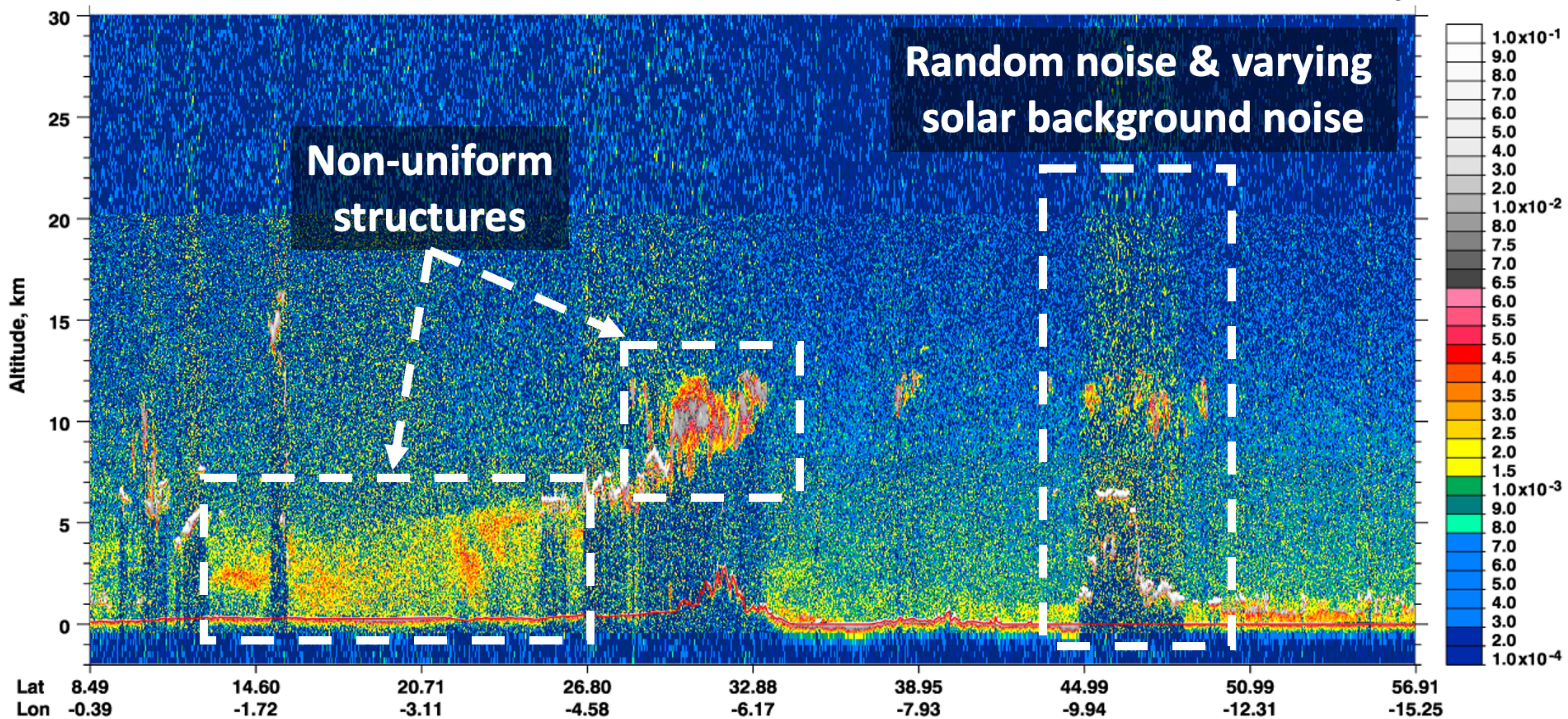
12/04/2023

Langley NASA CALIOP

532 nm Total Attenuated Backscatter, $\text{km}^{-1} \text{sr}^{-1}$ UTC: 2016-09-18 13:40:43.4 to 2016-09-18 13:54:12.1 Version: 4.10 Standard Daytime



532 nm Total Attenuated Backscatter, $\text{km}^{-1} \text{sr}^{-1}$ UTC: 2016-09-18 13:40:43.4 to 2016-09-18 13:54:12.1 Version: 4.10 Standard Daytime



Lat 8.49
Lon -0.39

Lat 14.60
Lon -1.72

Lat 20.71
Lon -3.11

Lat 26.80
Lon -4.58

Lat 32.88
Lon -6.17

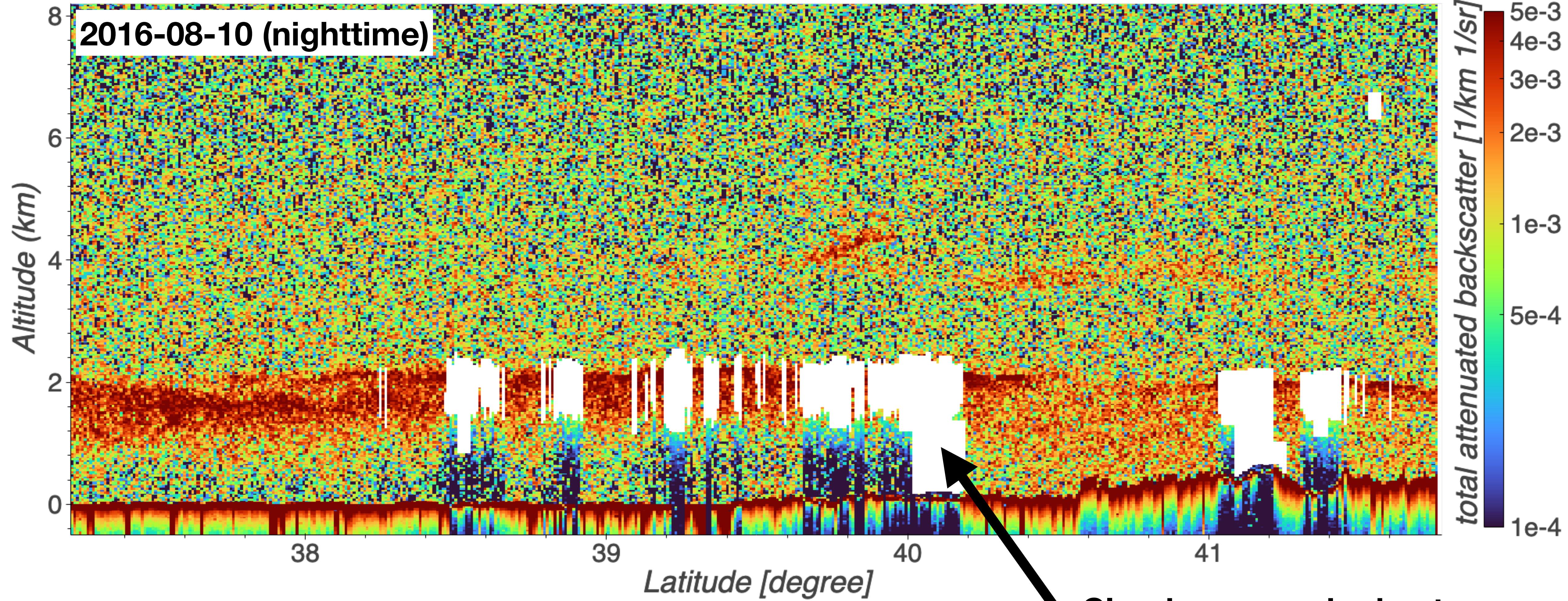
Lat 38.95
Lon -7.93

Lat 44.99
Lon -9.94

Lat 50.99
Lon -12.31

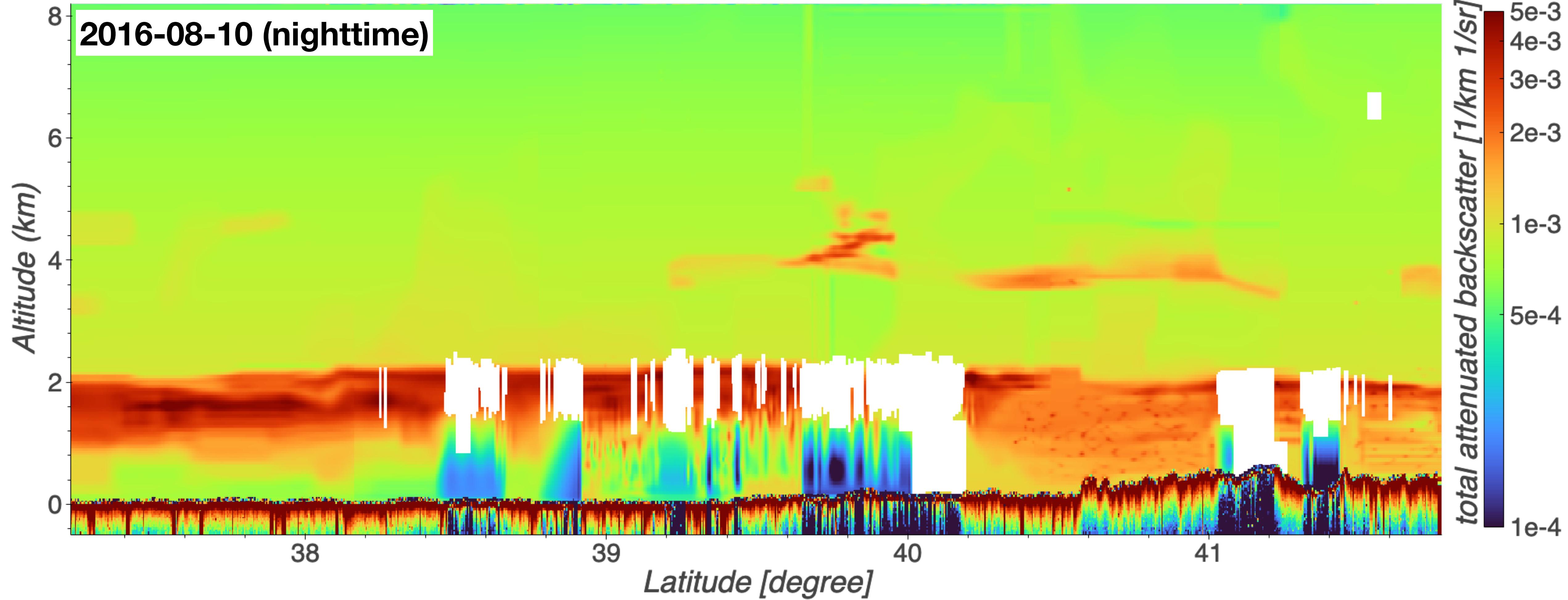
Lat 56.91
Lon -15.25

Noisy total attenuated backscatter (Horizontal 1km, Vertical 30m)



Clouds are masked out,
since we specifically interested
in denoising aerosols

Denoised total attenuated backscatter (Horizontal 1km, Vertical 60m)



The basic ingredients of denoising methods

- CALIOP noisy image: y
- Attenuated scattering ratio: x
- CALIOP forward model: $F(x)$
- CALIOP noise model: $\ell(y \mid F(x))$
- A priori assumption about
attenuated scattering ratio: $p(x)$

The basic ingredients of denoising methods

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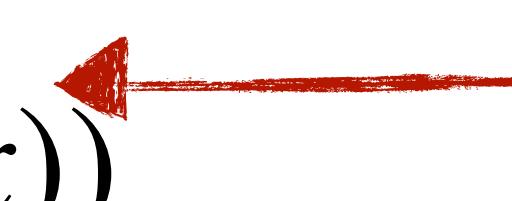
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 - Estimate parl. & perp. separately

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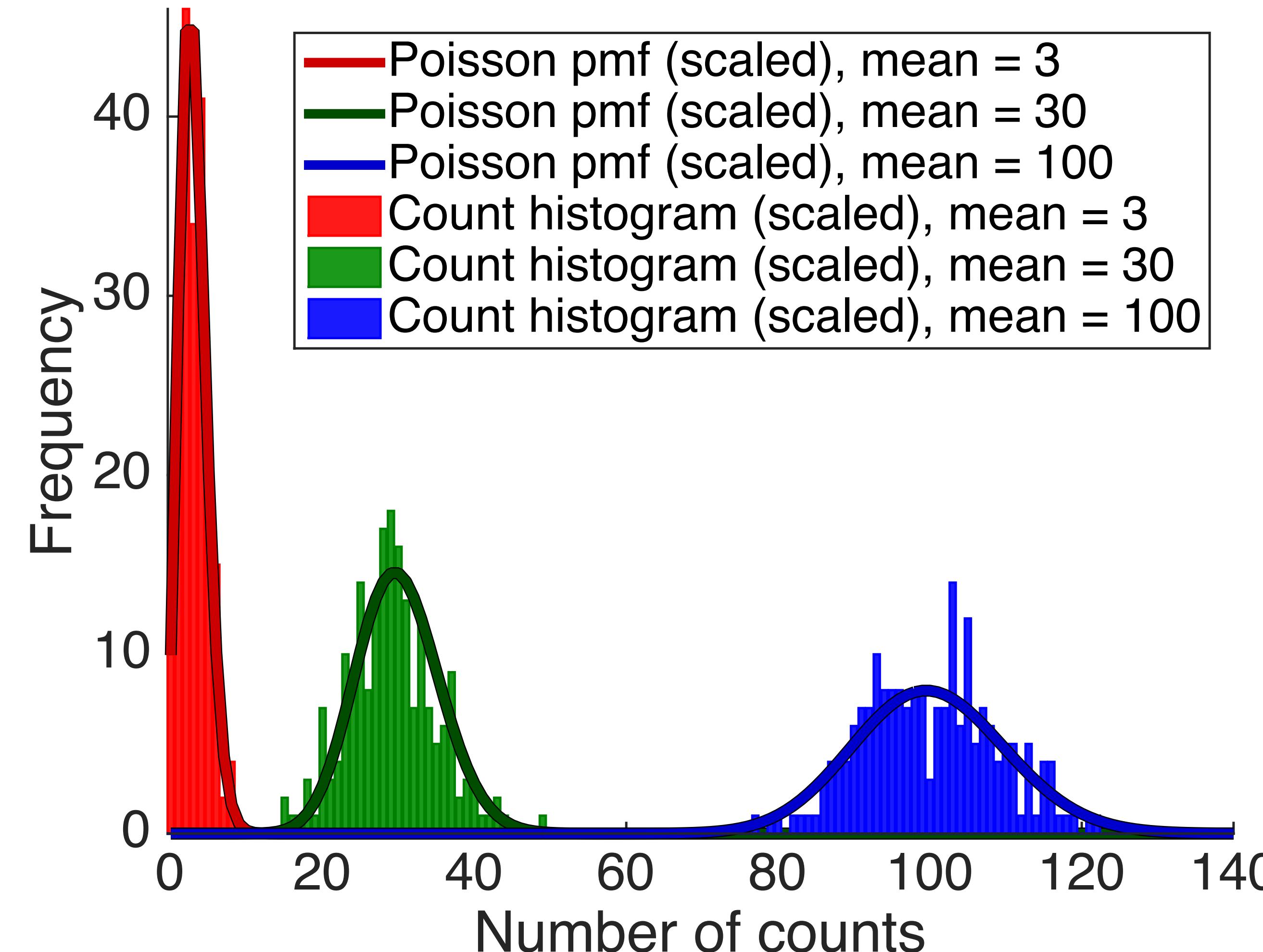
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 - Model expected value of y
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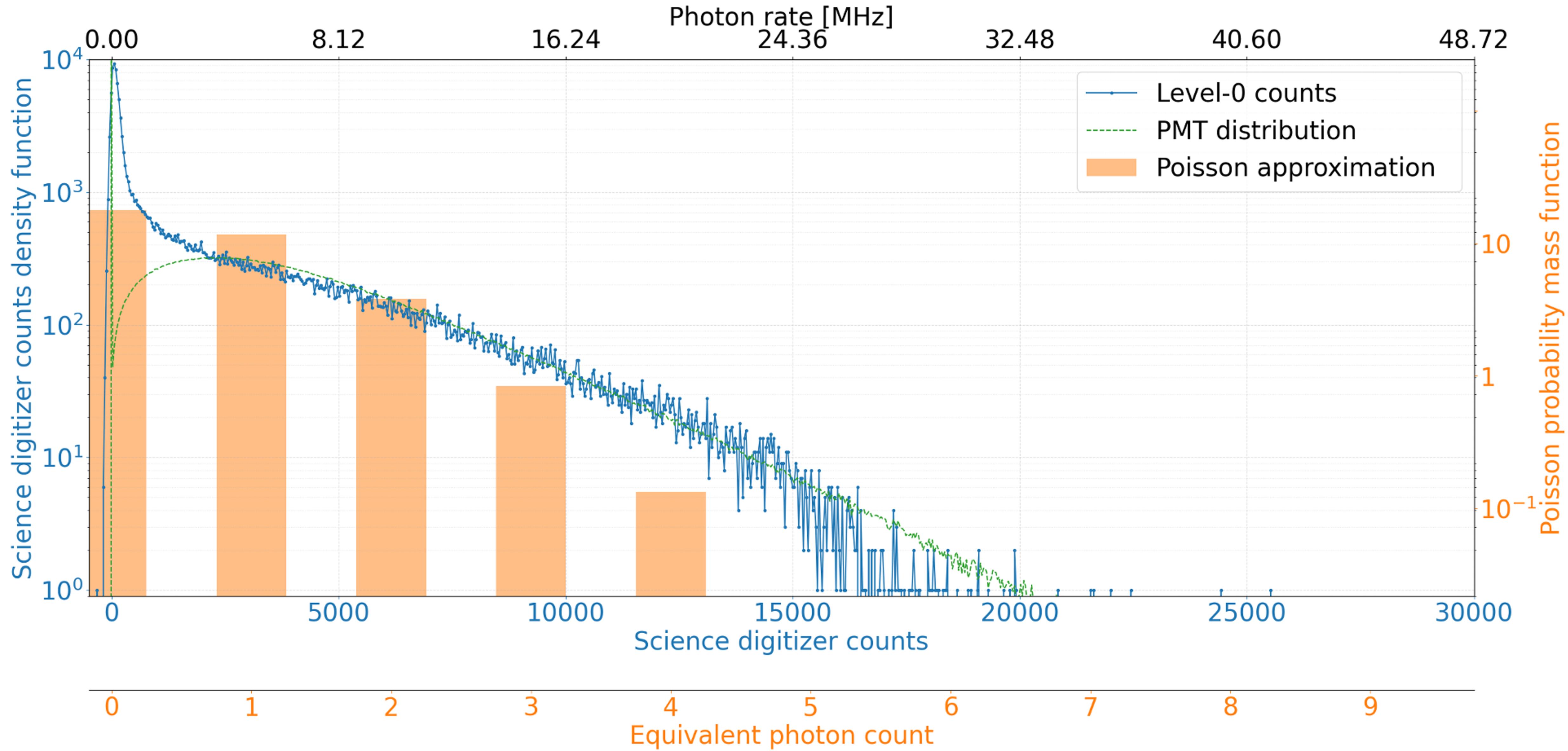
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 - Model noise statistical properties
- 

The noise model

Spatially-varying and signal-dependent noise variance



CALIOP noise probability distribution



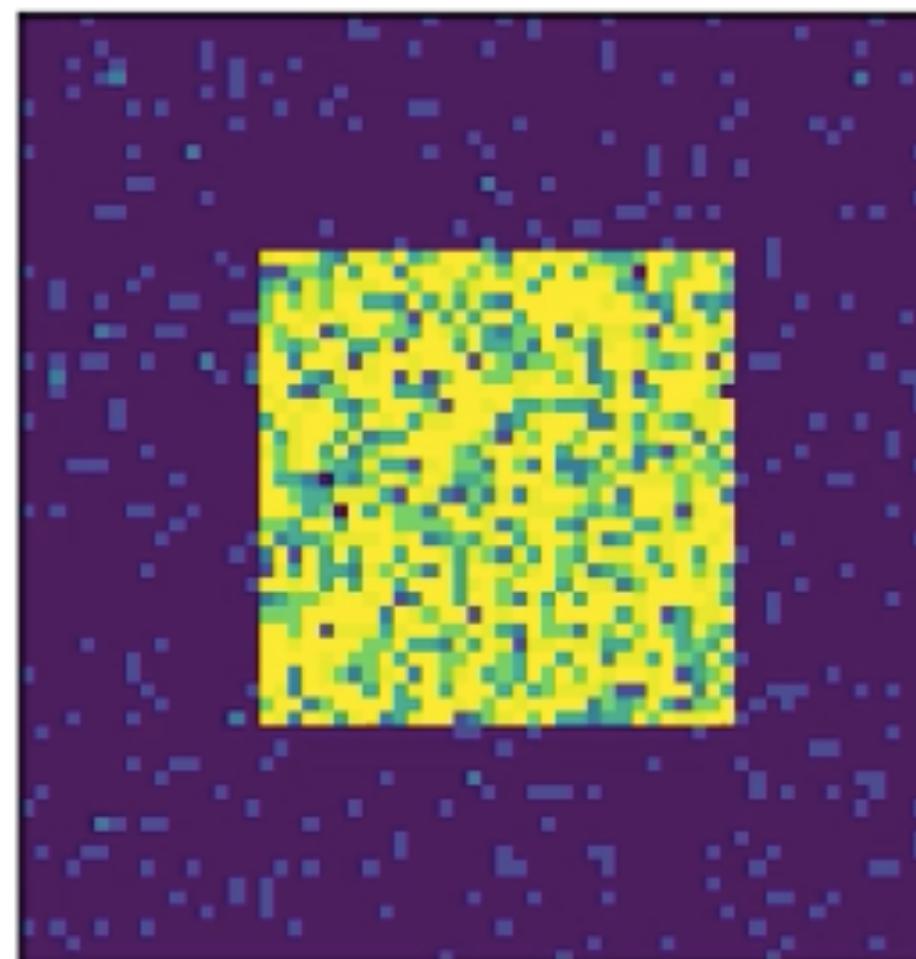
Starting with something familiar

The formulation of optimal estimation

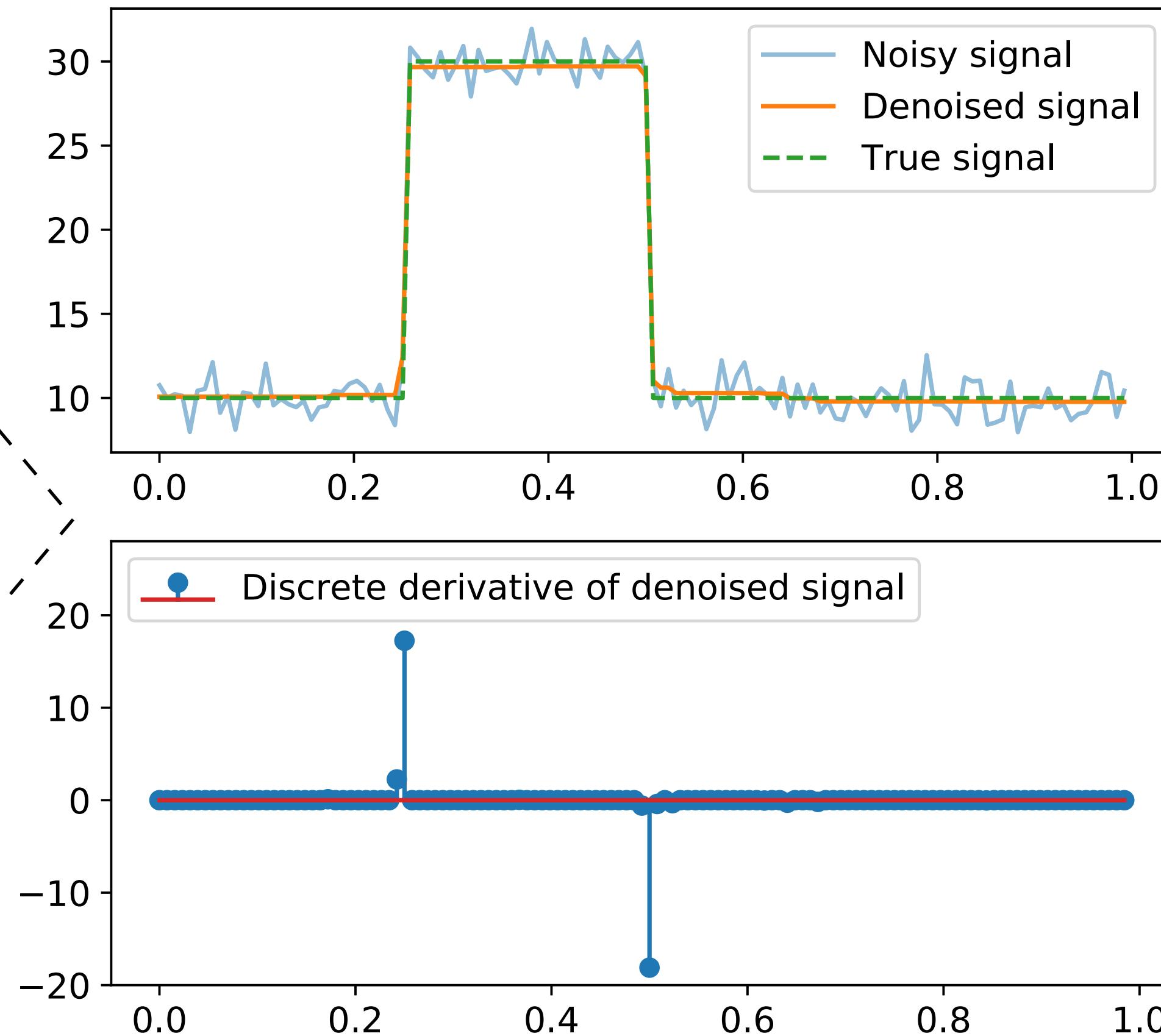
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 - Model noise statistical properties
 - Promote structure / spatial + temporal correlation in image
- 

Poisson total variation (PTV)

PTV approximates the image as piecewise constant

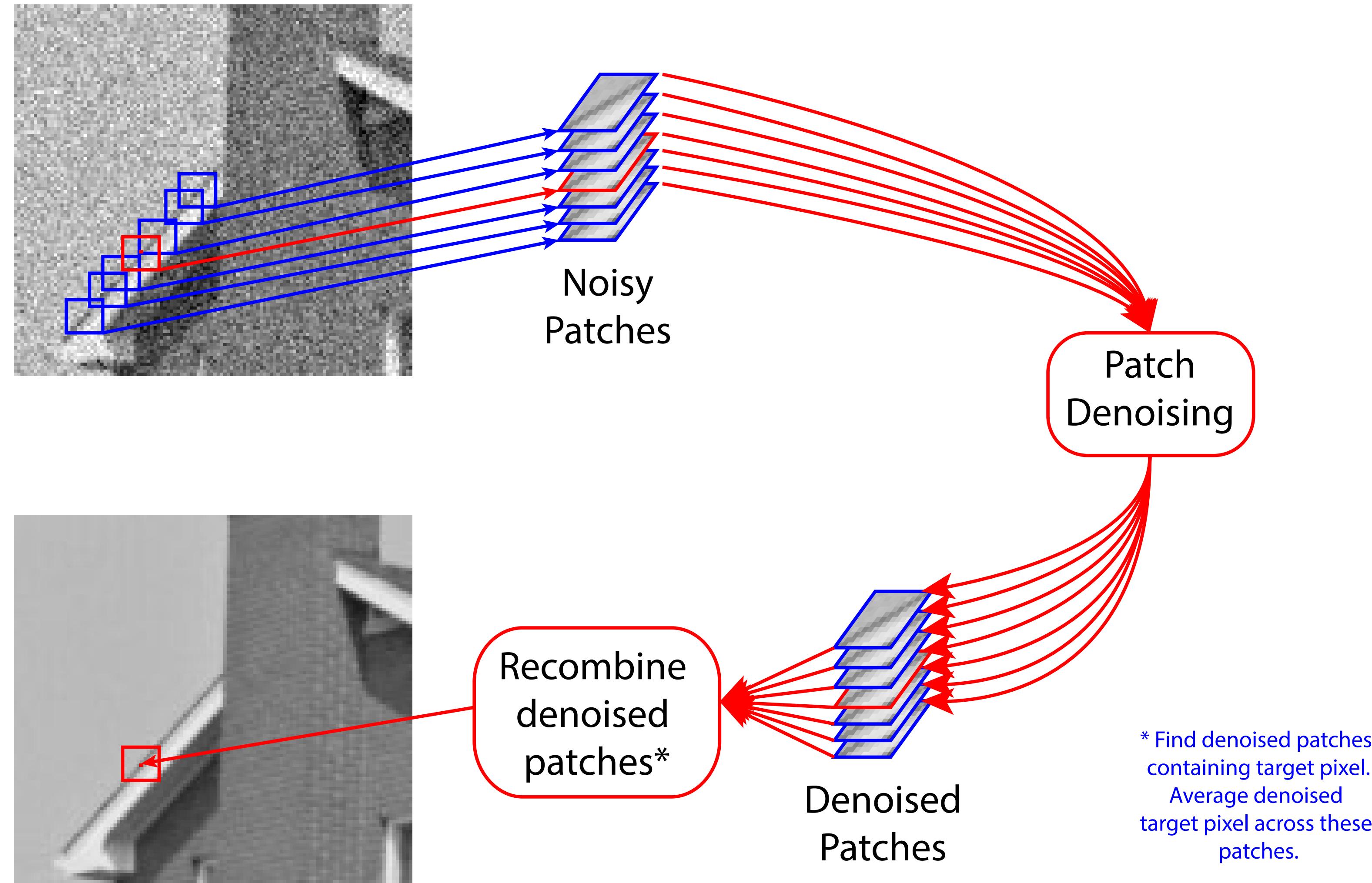


cross-section

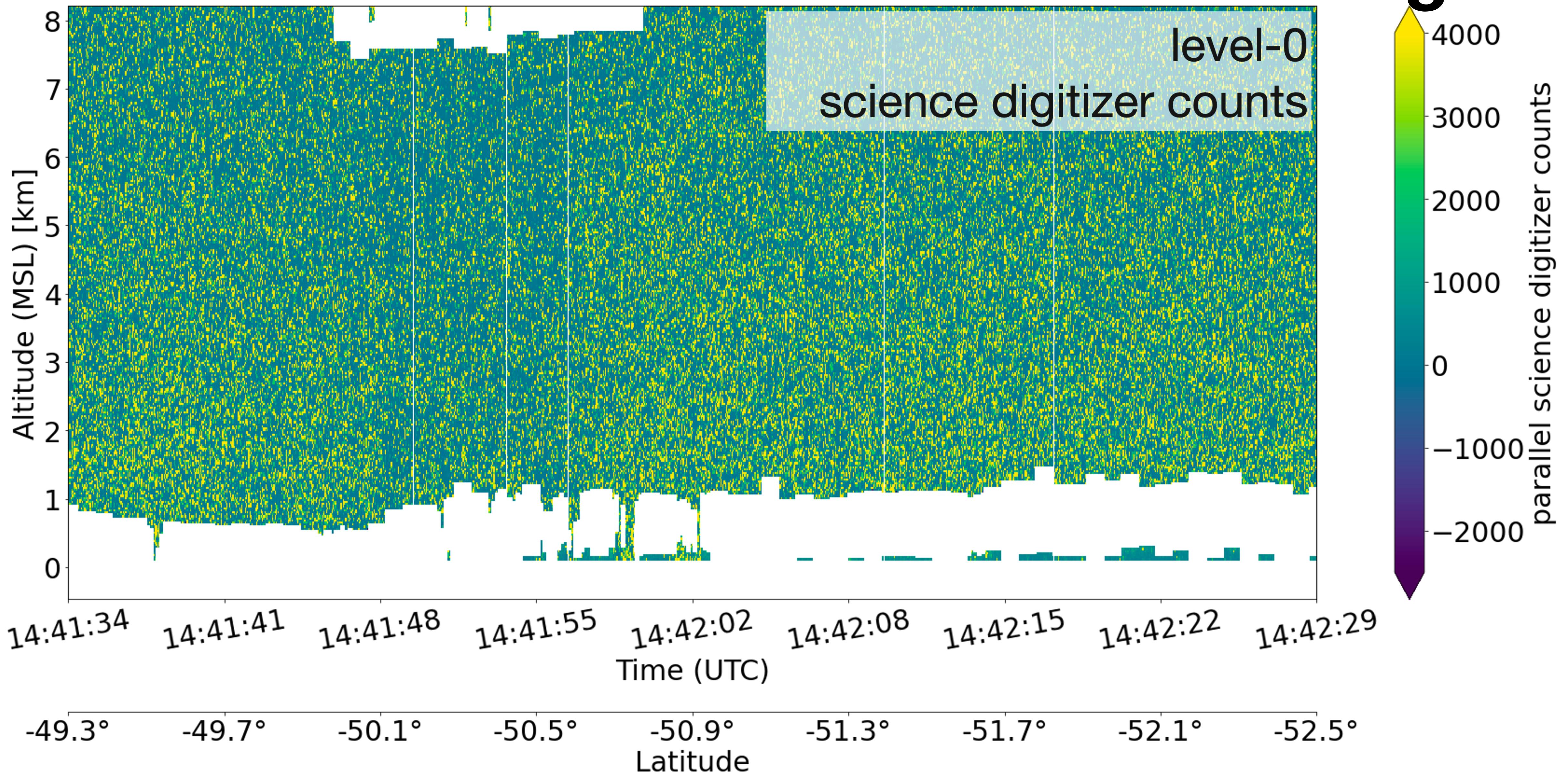


Patch based denoising

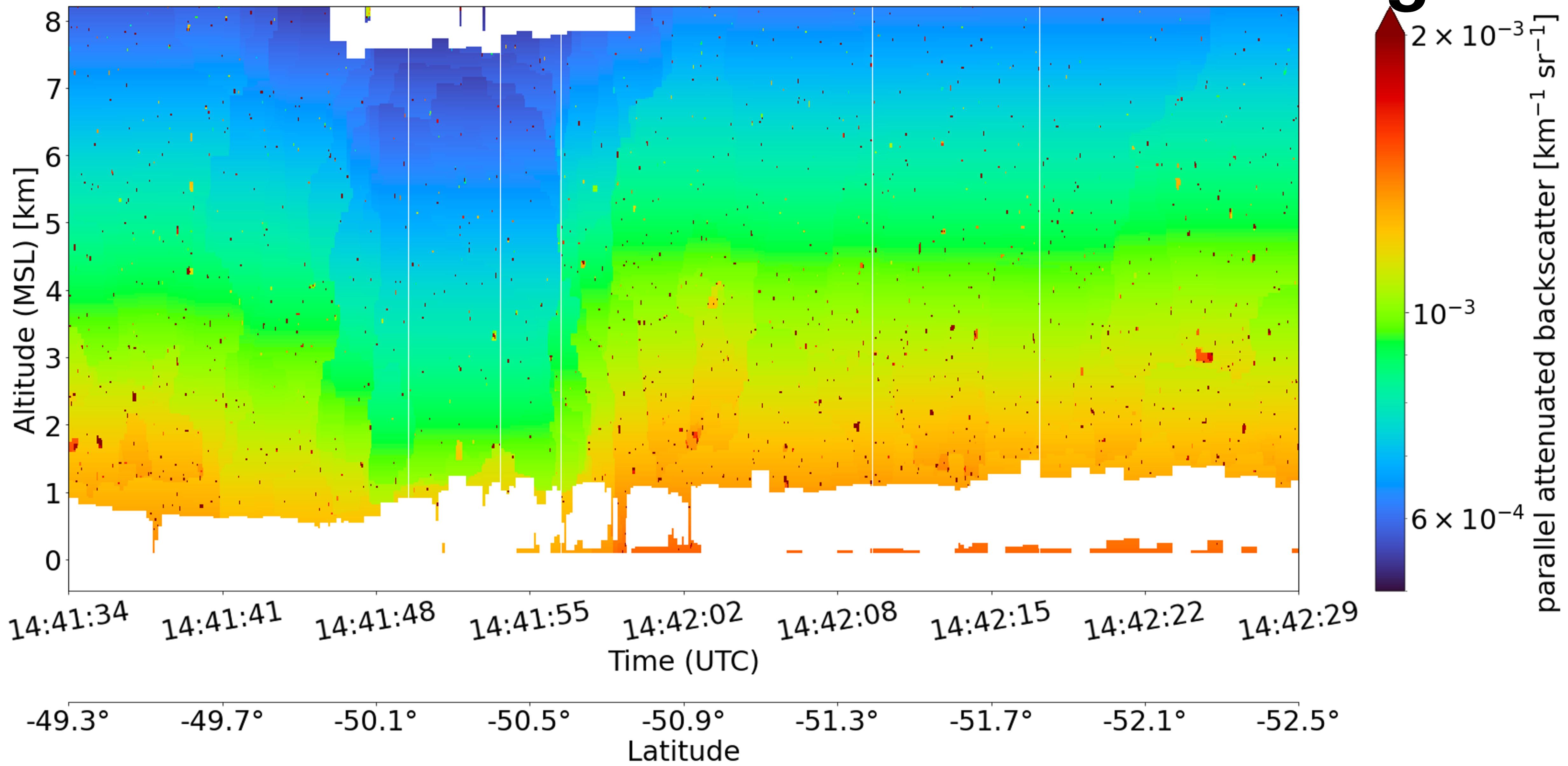
Exploit redundancy in image that allows for accurate approximation of a richer class of images



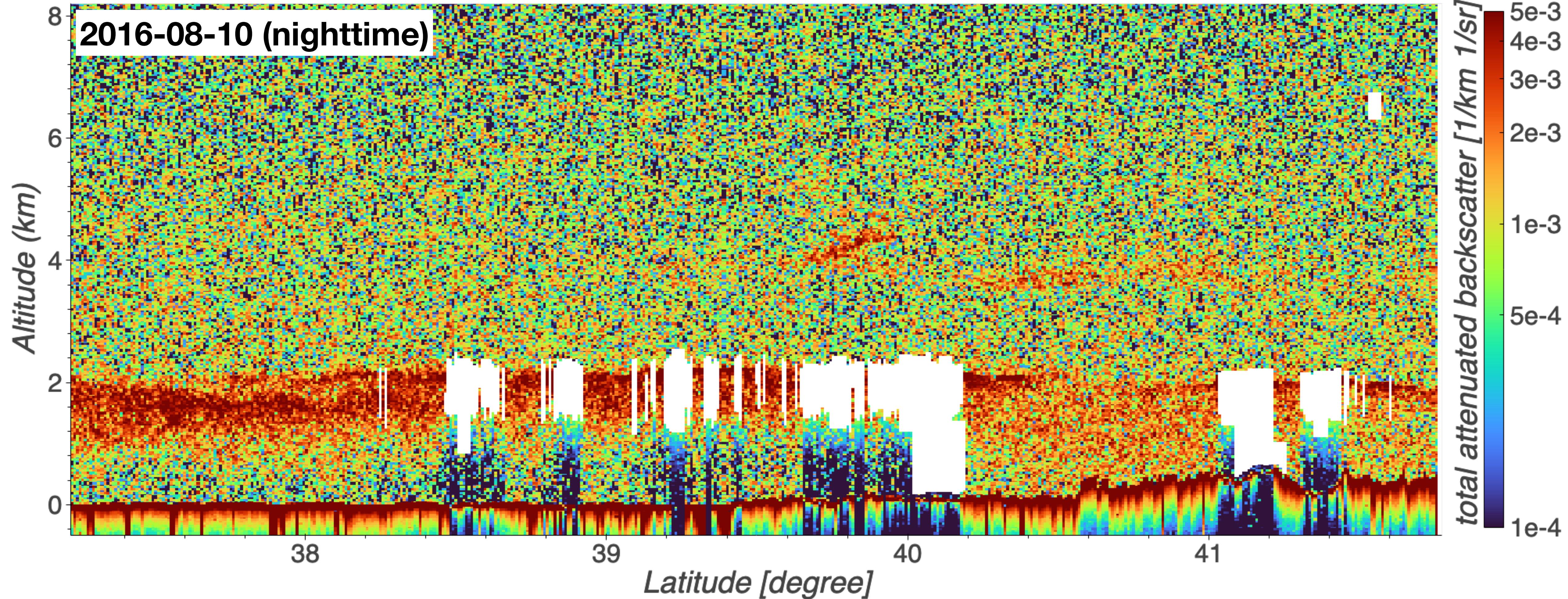
@ 2022 CALIPSO science team meeting



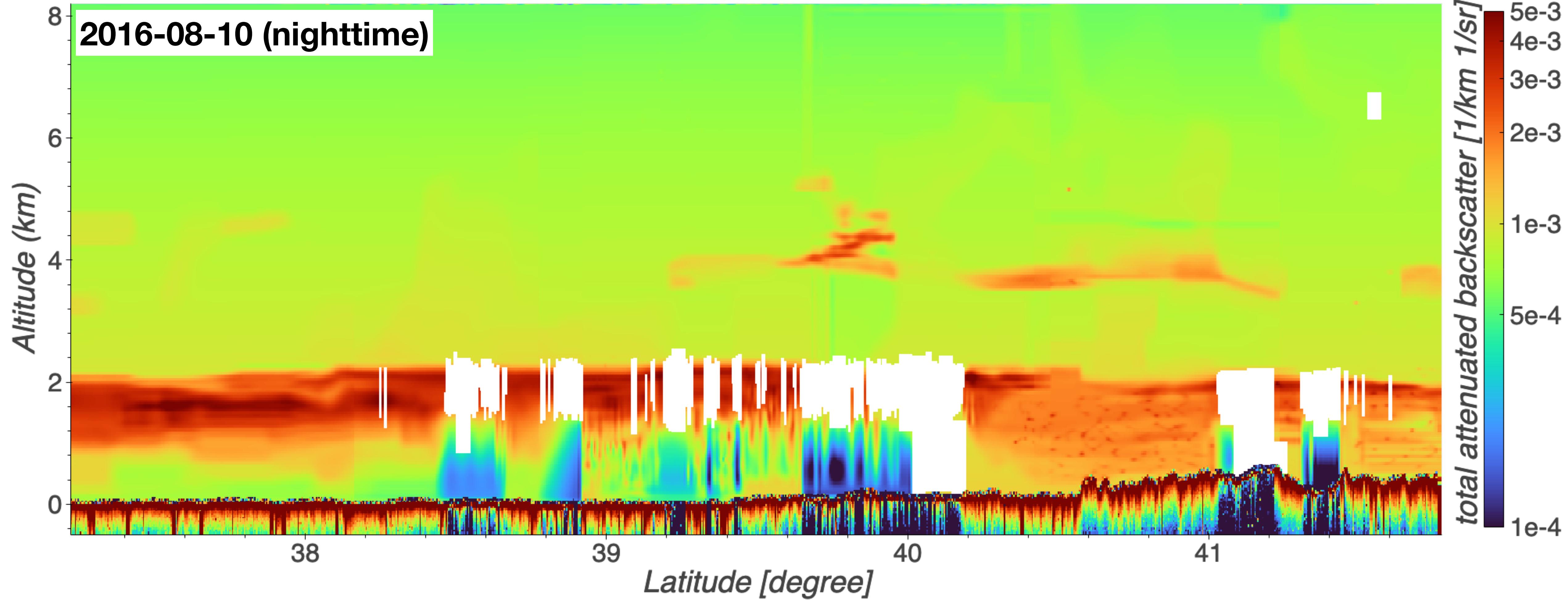
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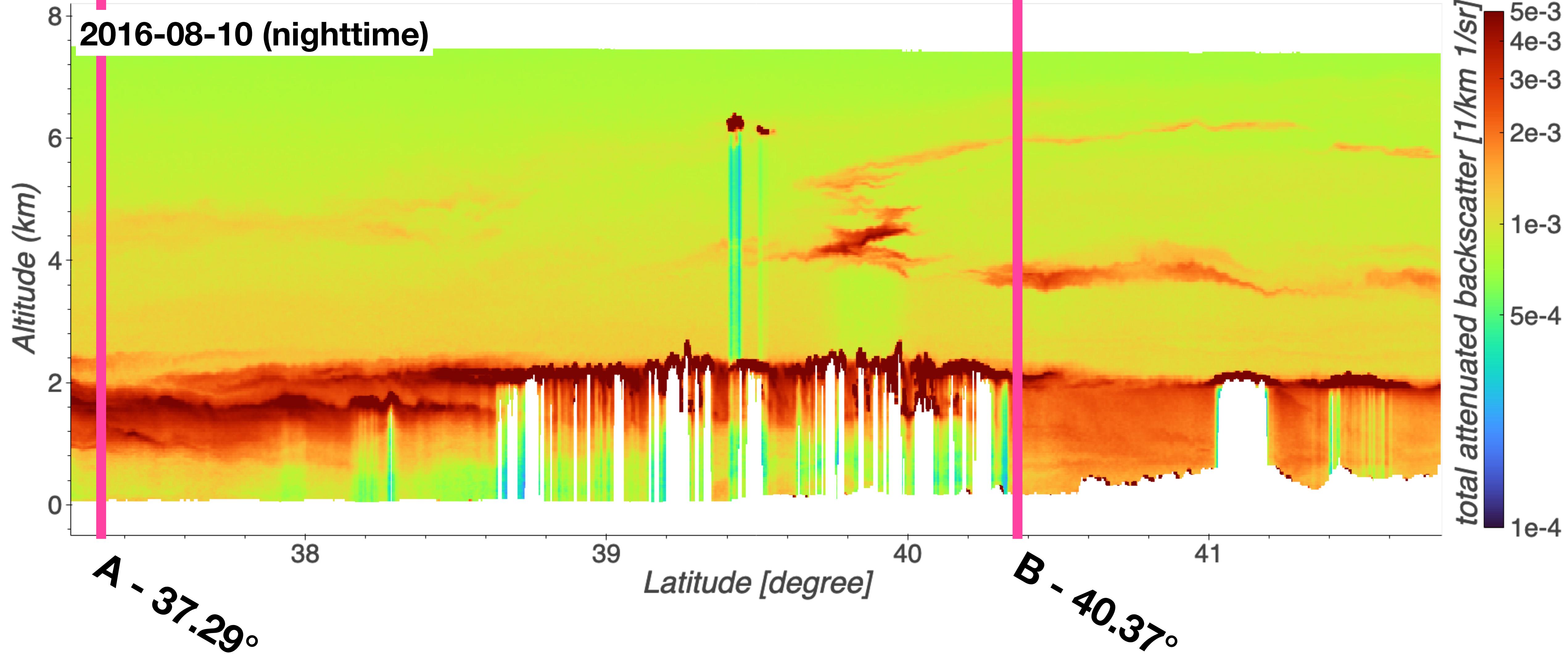
Noisy total attenuated backscatter (Horizontal 1km, Vertical 30m)



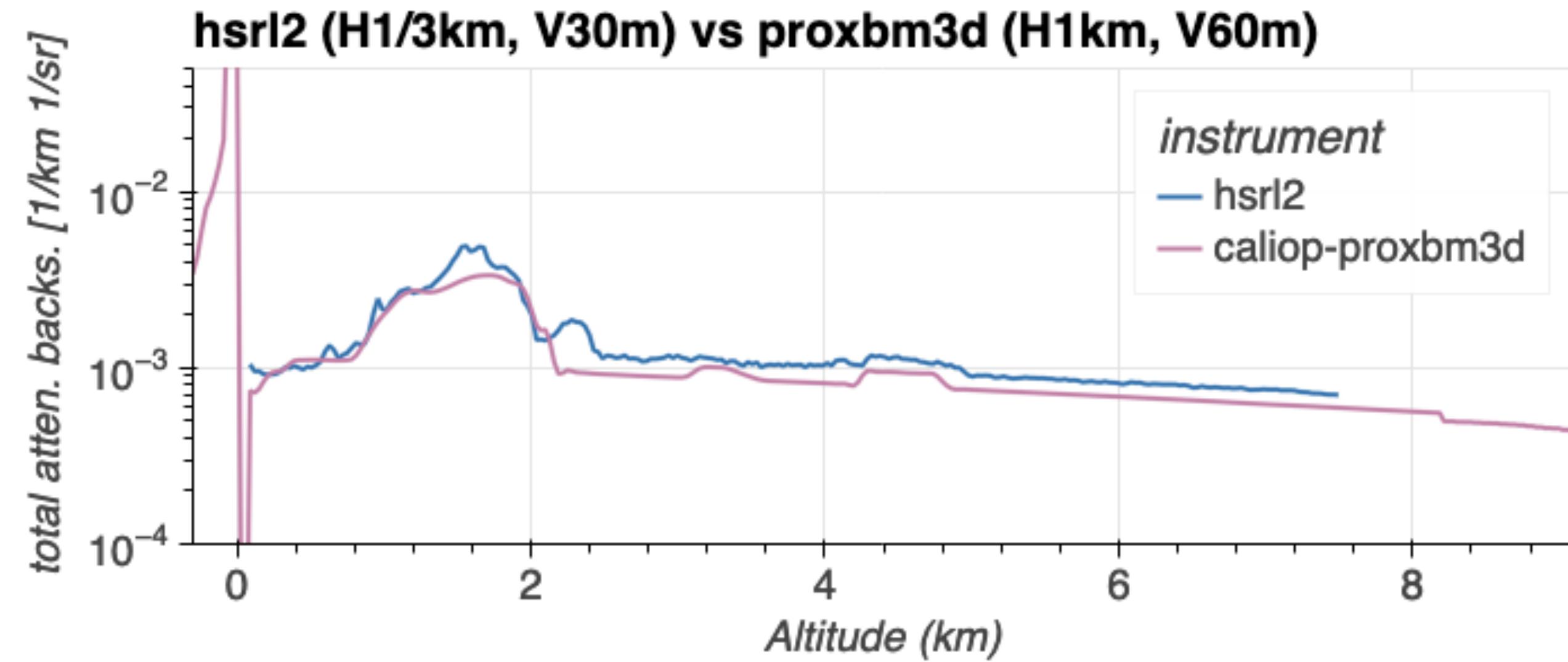
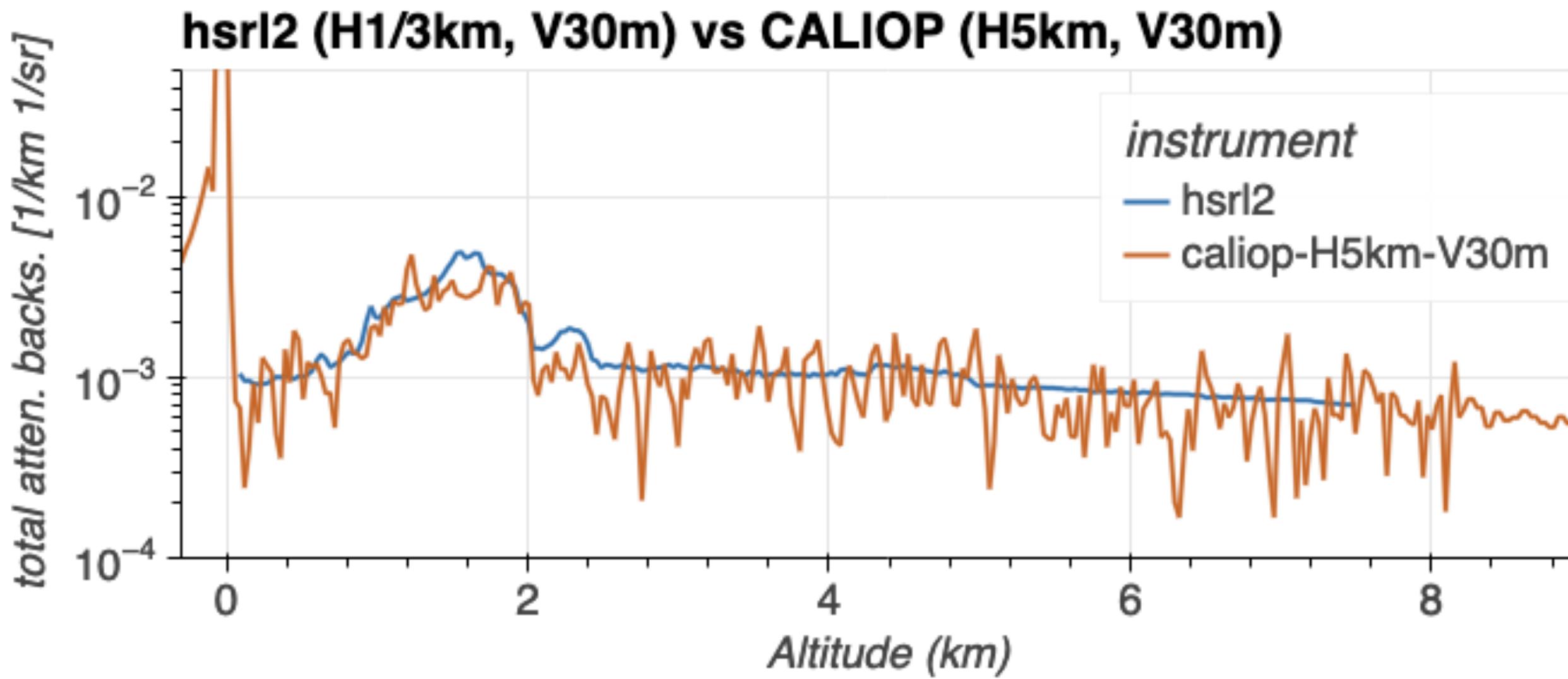
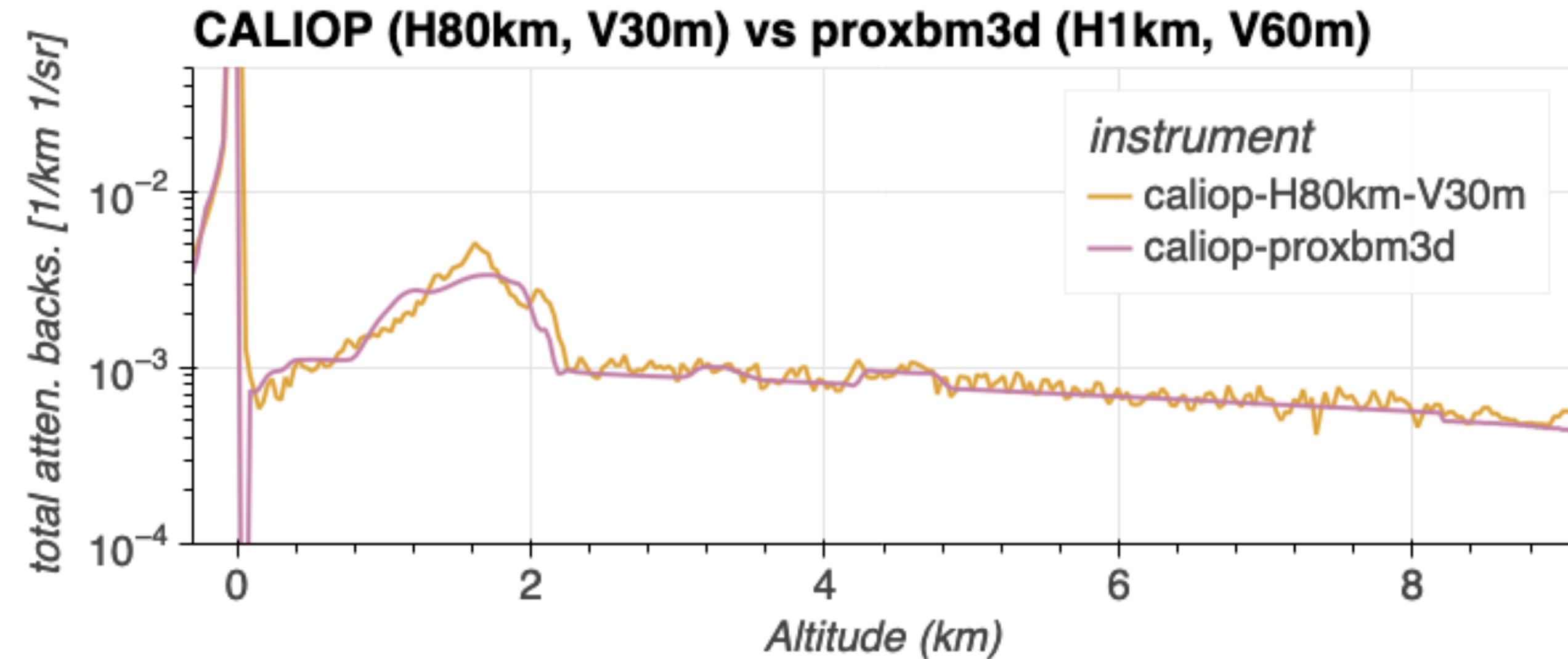
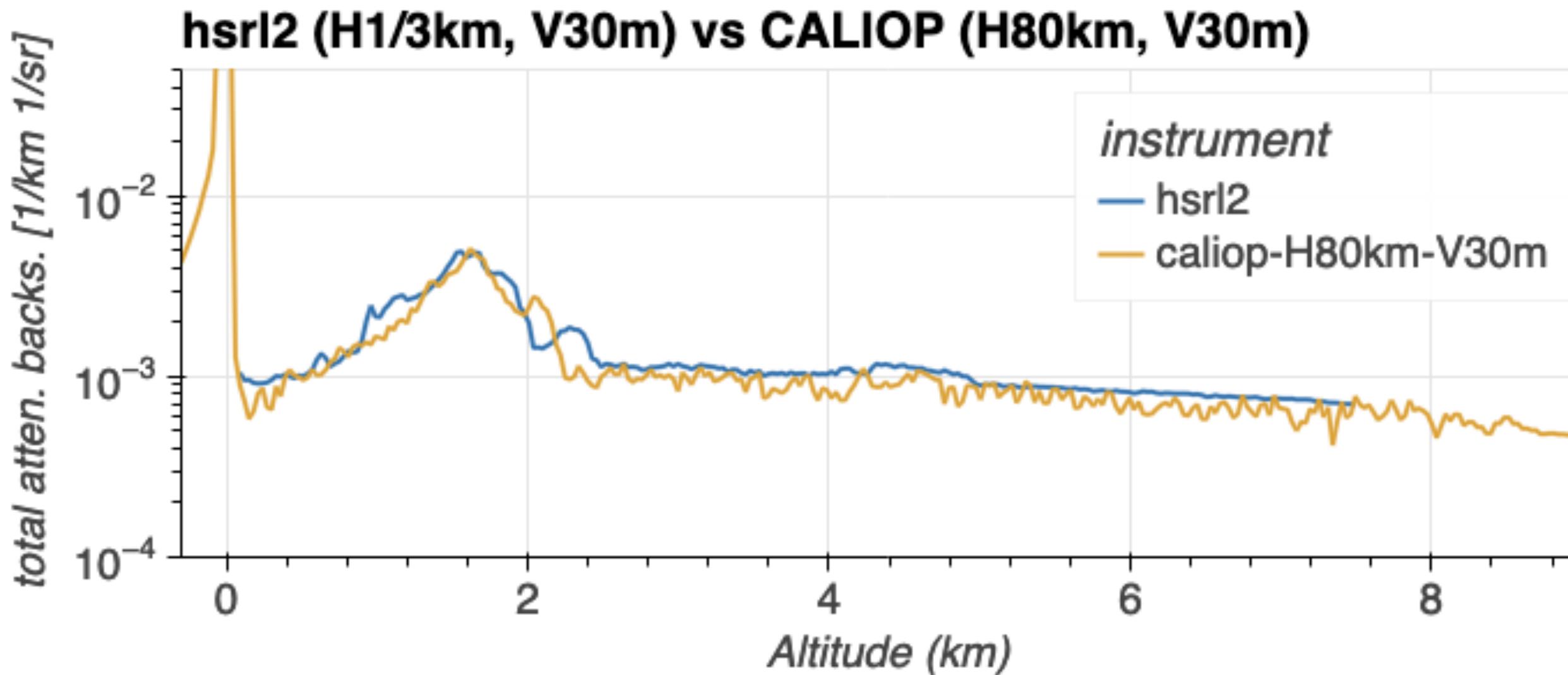
Denoised total attenuated backscatter (Horizontal 1km, Vertical 60m)



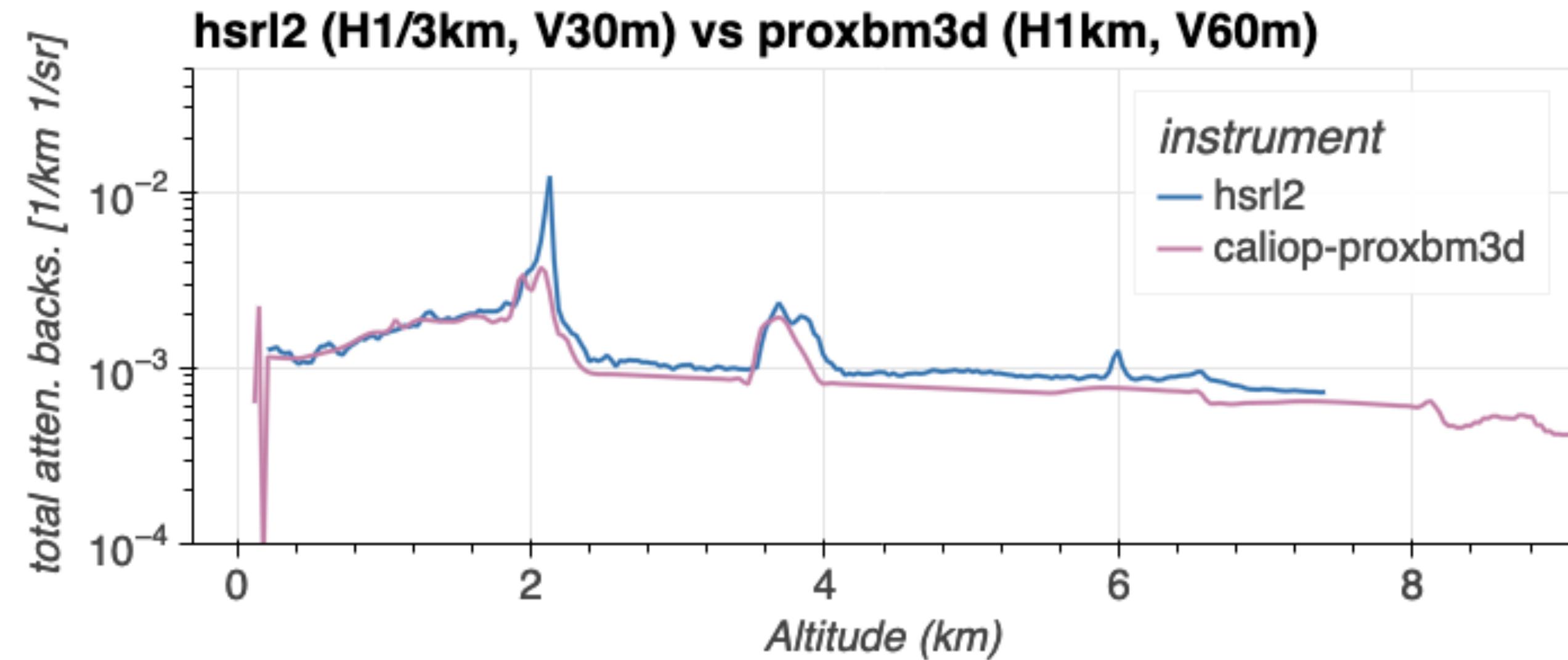
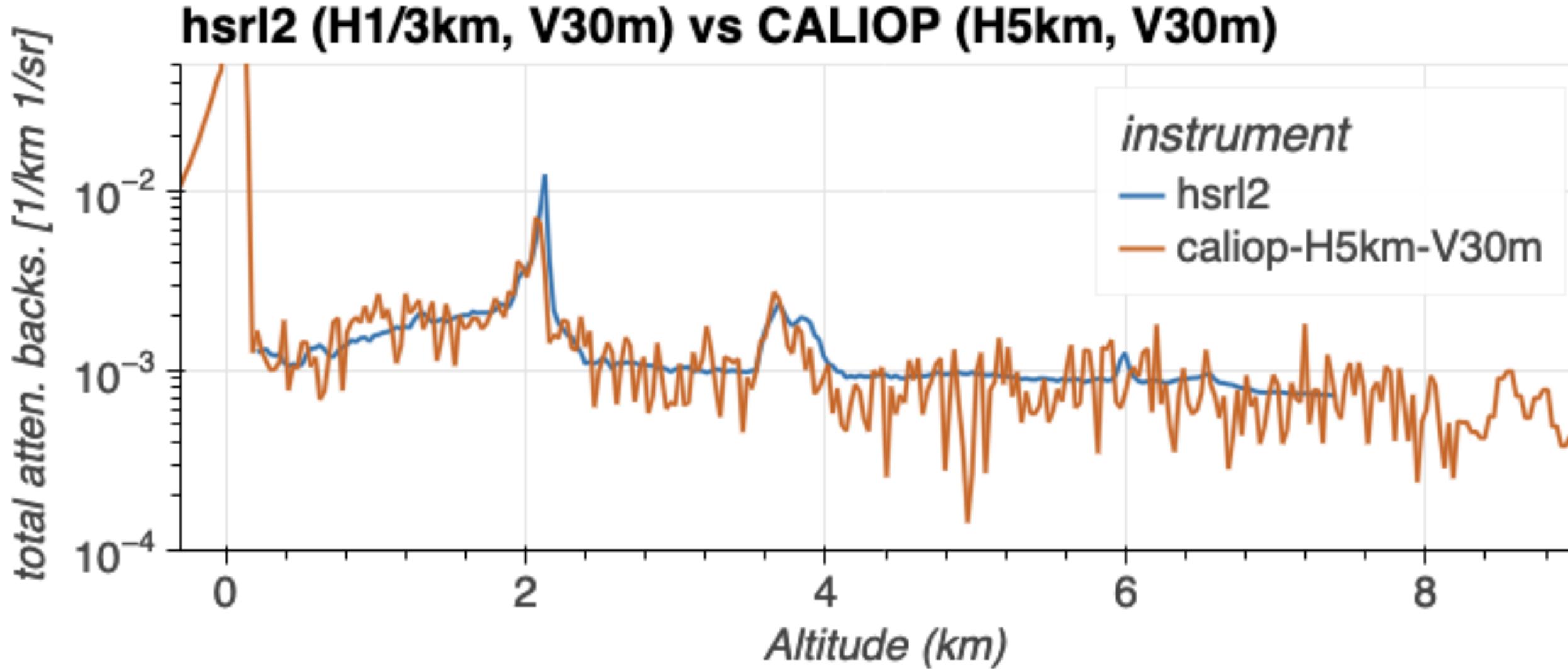
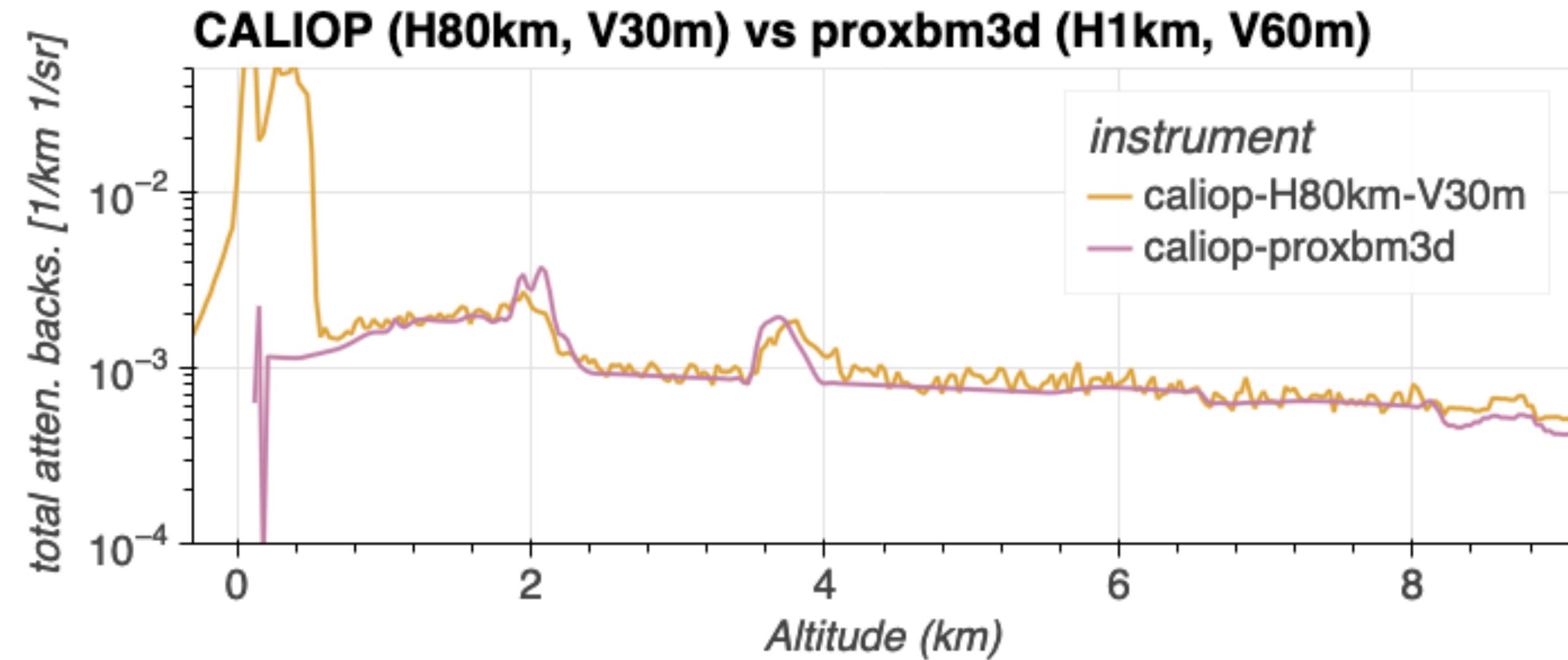
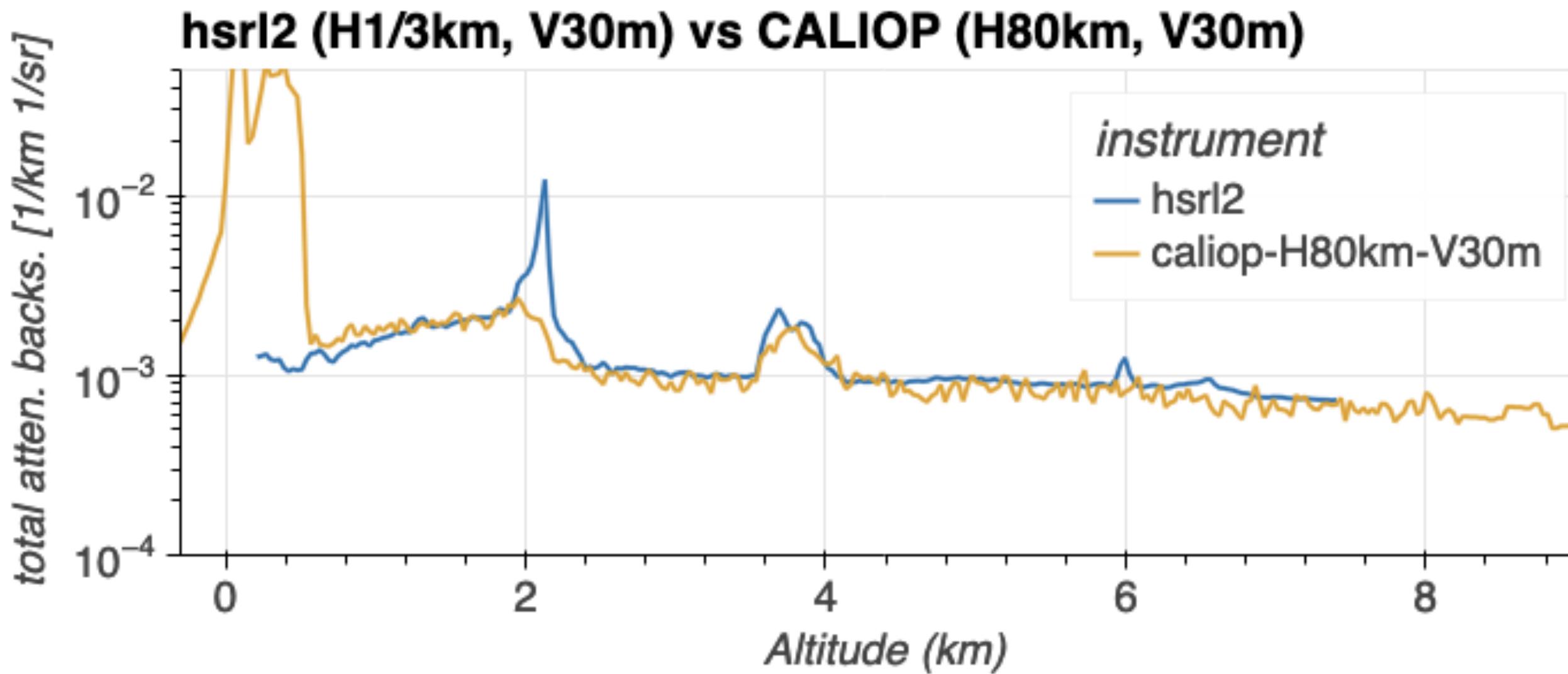
HSRL2 total attenuated backscatter (Horizontal 1/3km, Vertical 30m)



Profile A - 37.29°



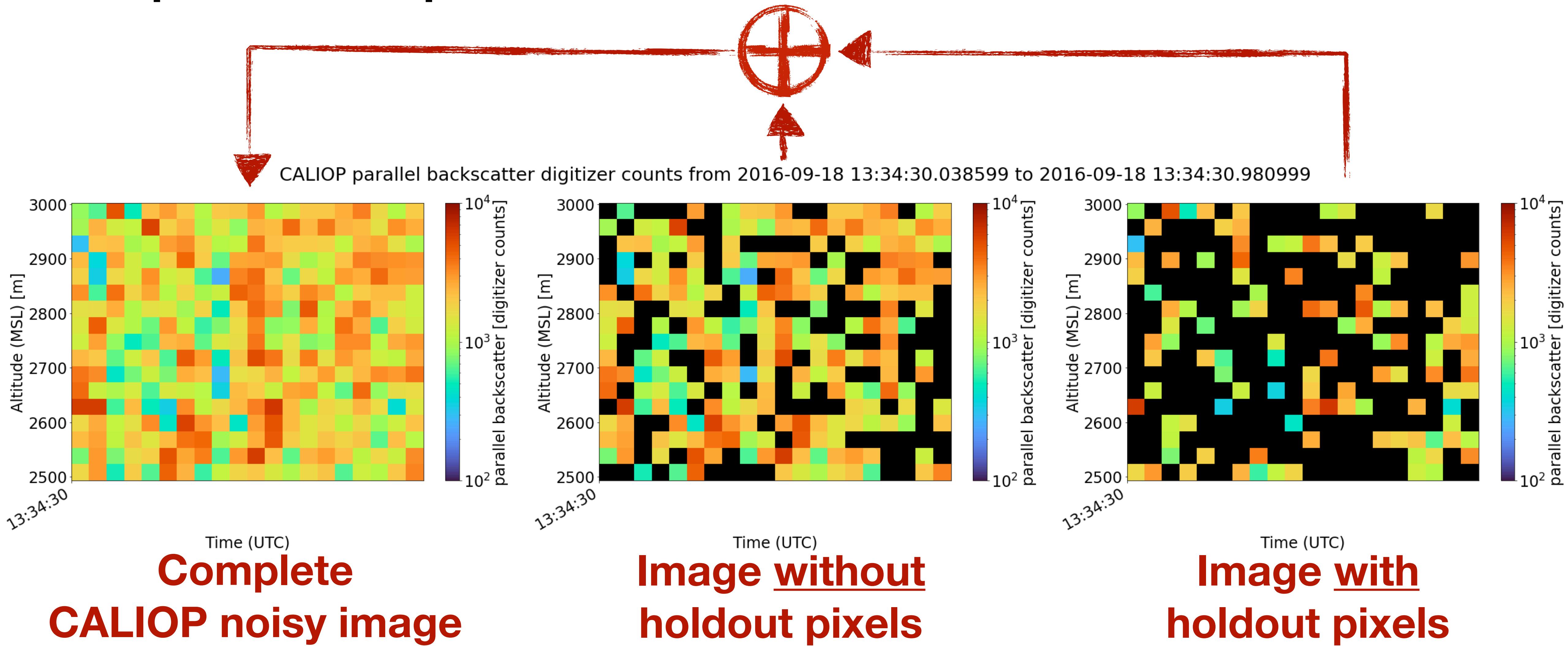
Profile B - 40.37°



Backup slides

Cross-validation: Choosing the regularization parameter

Step 1: Holdout pixels



Cross-validation: Choosing the regularization parameter

Step 2: Denoise and interpolate over holdout pixels

1) For regularization parameter λ denoise & interpolate

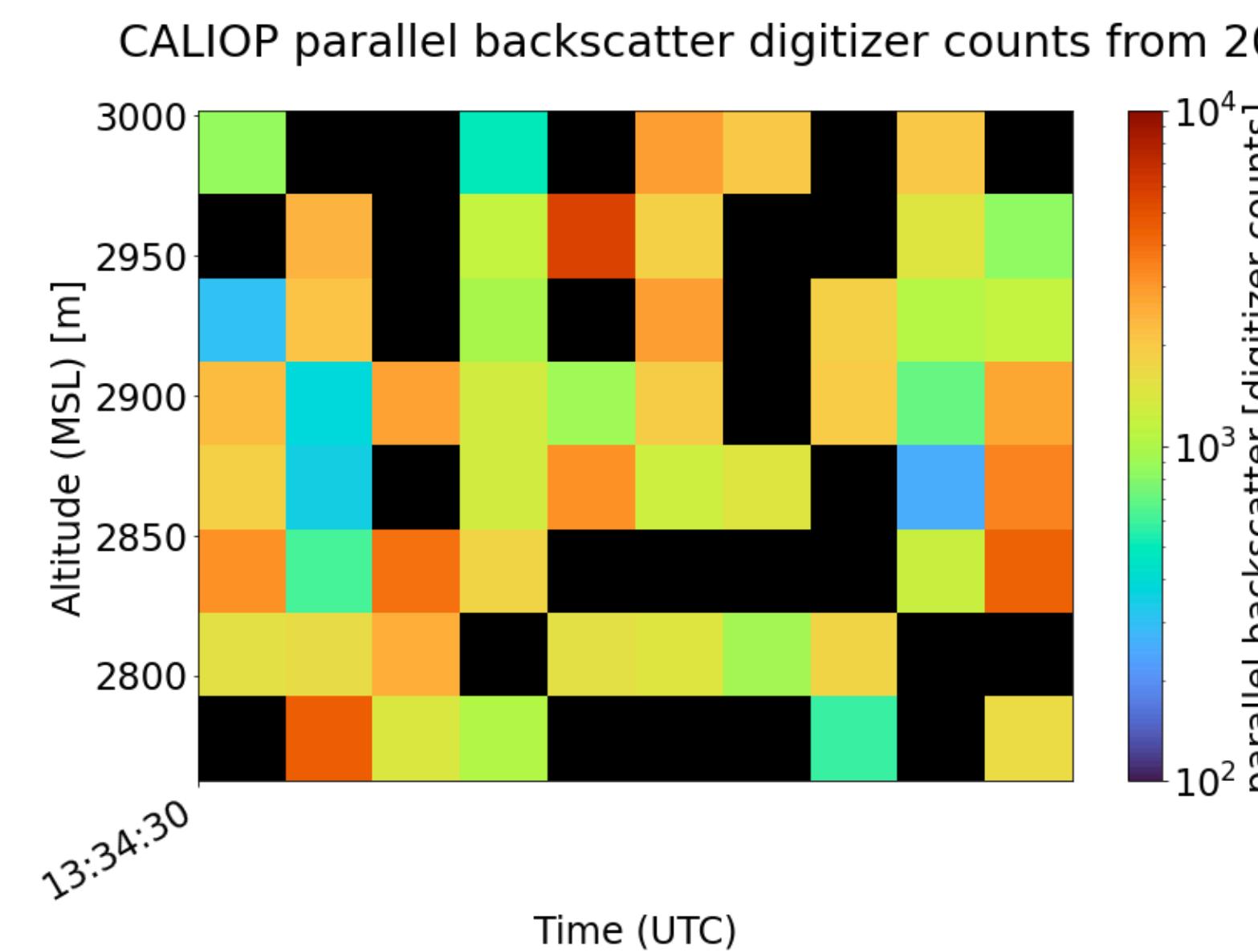


Image without holdout pixels

2) Choose estimate with regularization parameter λ which best fits holdout pixels

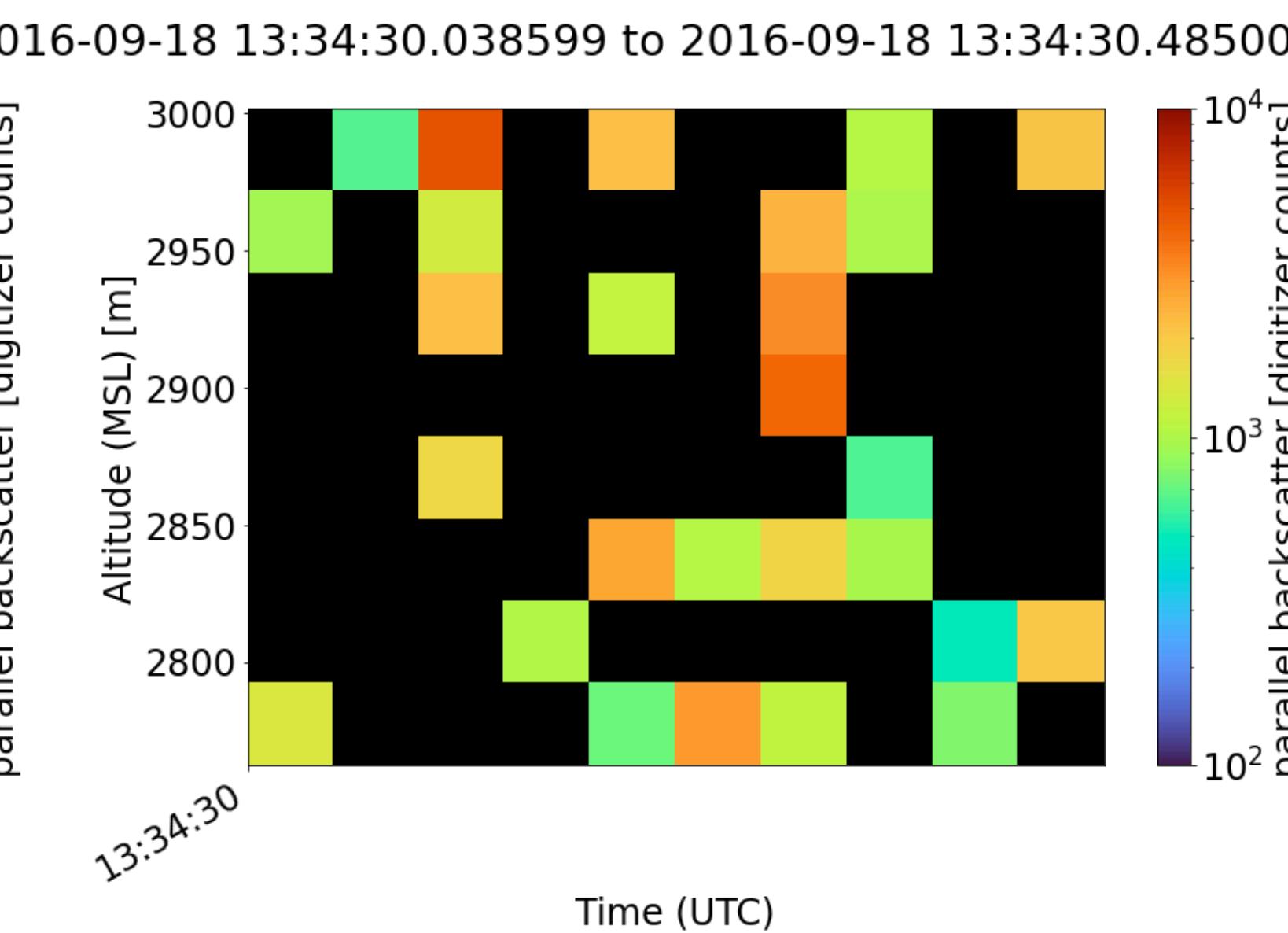


Image with holdout pixels

The three key ideas that OE shares with regularized maximum likelihood estimation

1) Noise model quantifies
goodness of fit
between $F(x)$ and y



$$l(y \mid F(x)) + \lambda p(x)$$

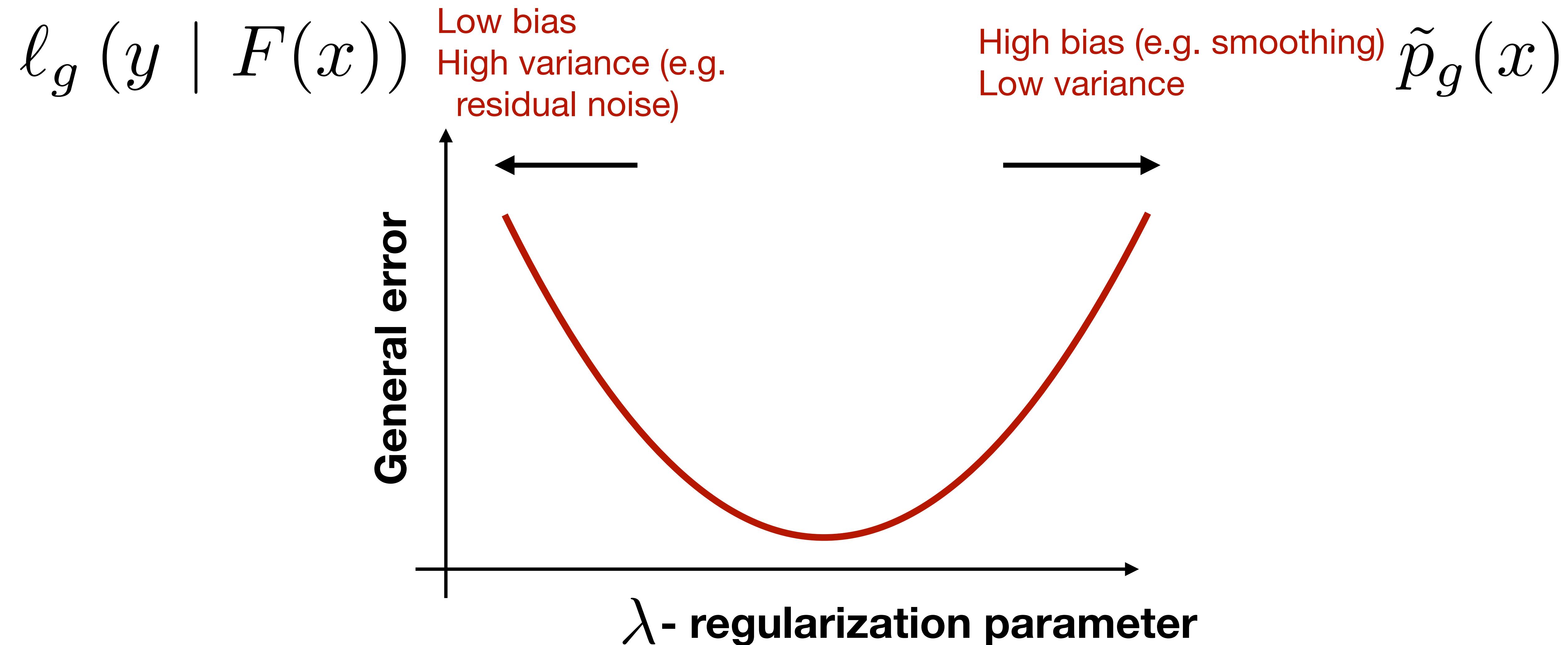
2) Regularizer function
that promotes a
priori about x



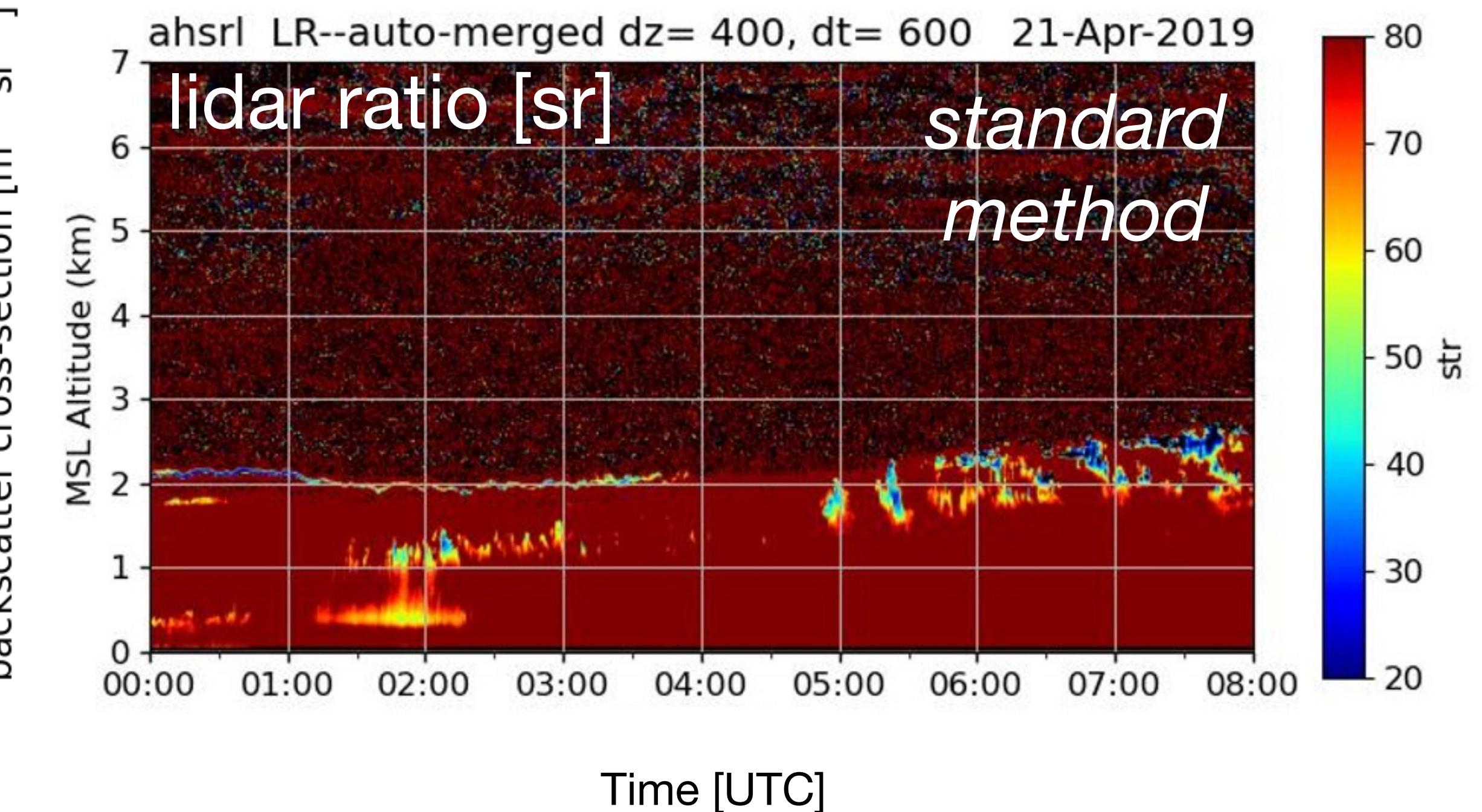
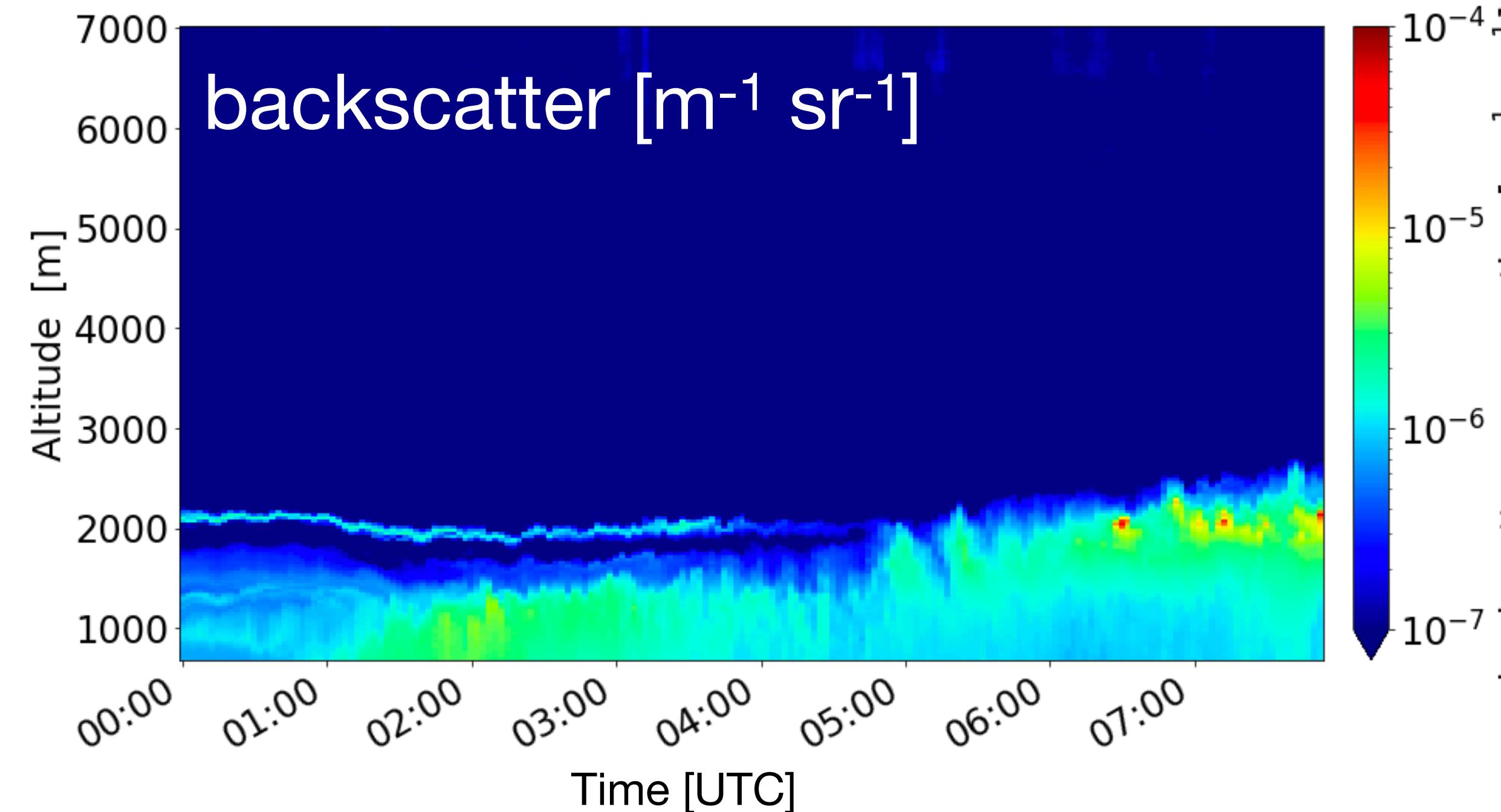
3) Regularization parameter sets the
degree to which the
a priori of x is promoted

Error vs the regularization parameter

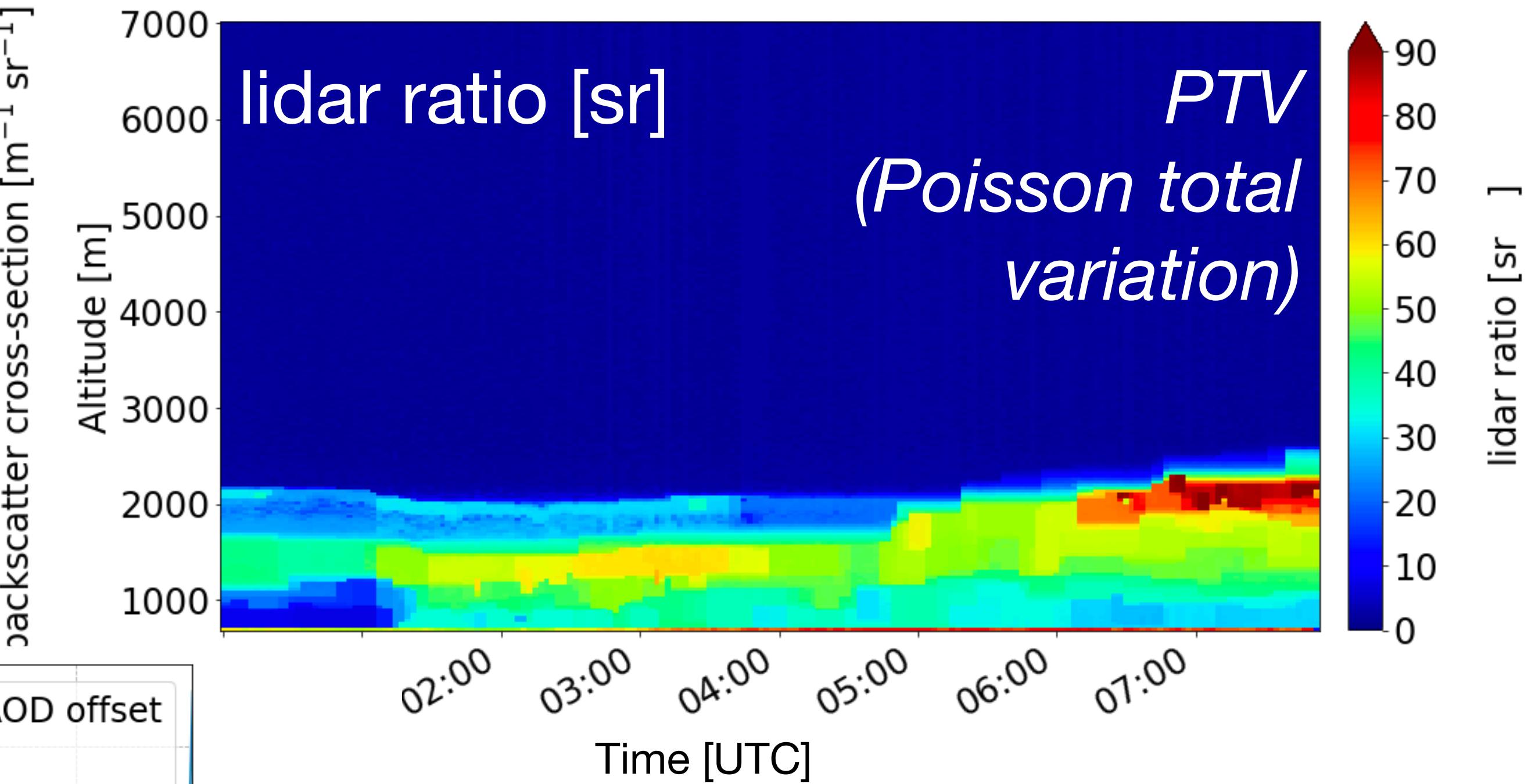
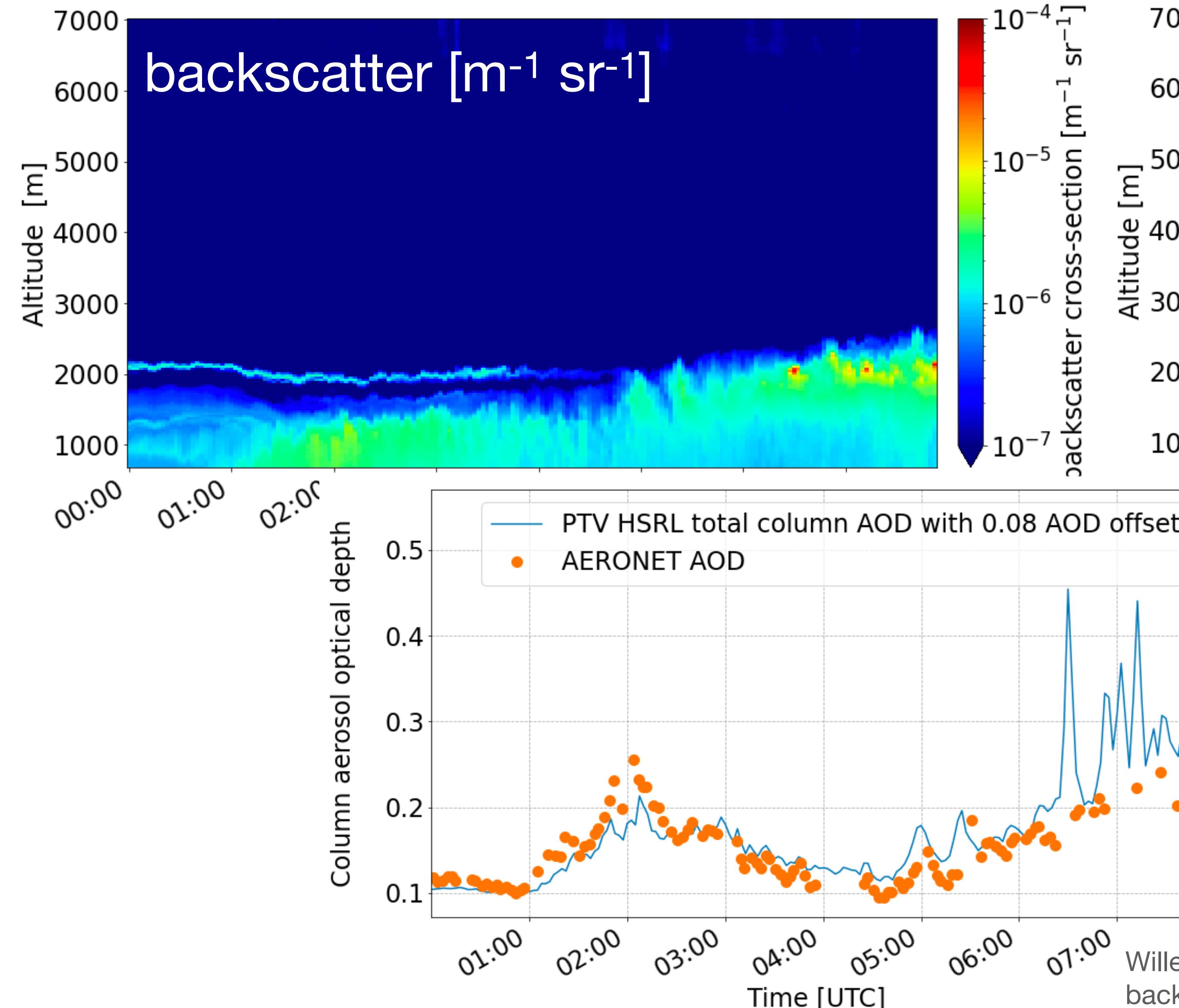
Intuition behind the regularization parameter



Denoising UW High Spectral Resolution Lidar data



Denoising UW High Spectral Resolution Lidar data



Denoising NCAR Micro Pulse DIAL (MPD) data

